

Techni-Dilaton as Dark Matter

Deog Ki Hong

Pusan National University

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with K. Y. Choi and S. Matsuzaki

Introduction

Review of WTC

Very light techni-dilaton

Dark matter TD

Conclusion

Introduction

- ▶ Higgs holds a key to BSM, since it is sensitive to short distance physics.
- ▶ But, we have not found Higgs yet. Current mass bound is

$$114 \text{ GeV} \lesssim m_H \lesssim 157 \text{ GeV}$$

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Introduction

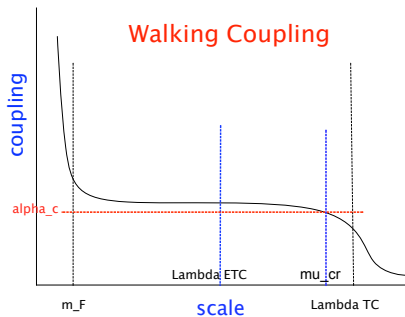
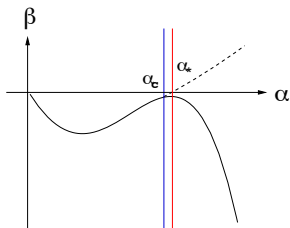
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Introduction

- ▶ Modern TC is called “Walking Technicolor” (Holdom '81, Yamawaki et al '86, Appelquist et al '86)



Introduction

- ▶ As a candidate for physics BSM, it will be nice if TC explains dark matter as well.
- ▶ Indeed, I will show that WTC can have very light dilaton, techni-dilaton as Nambu-Goldstone bosons, associated with spontaneously broken (approximate) scale symmetry, which can be a good candidate for DM. (Cf. Techni-baryons)

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Review of WTC

- ▶ Introduce new strong dynamics in addition to SM, which is confined and breaks chiral symmetry:

$$SU(3)_c \times SU(2)_L \times U(1)_Y \times G_{TC}$$

- ▶ Introduce new particles, N_F techniquarks, which transform as

$$Q_L^{TC} \sim (3, 2, y_L, r), \quad Q_R^{TC} \sim (3, 2, y_R, r)$$

- ▶ such that theory is anomaly-free, (asymptotically free) and has a (quasi) IR fixed point.

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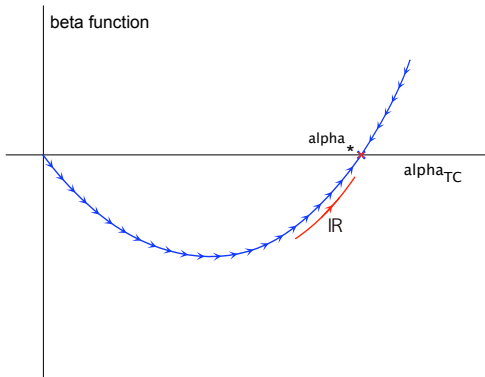
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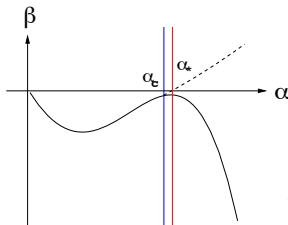
- ▶ TC has $SU(N_F)_L \times SU(N_F)_R$ chiral symmetry. (We may need techni-leptons to be anomaly-free.)

Review of WTC

- ▶ We assume that chiral symmetry is spontaneously broken by TC interactions: the critical coupling for χ SB, $\alpha_c < \alpha_*$.

$$\alpha_c \approx \frac{\pi}{3C_2(r)}$$

- ▶ We assume that $\alpha_c \approx \alpha_*$ to have walking behavior.
- ▶ Once techni-fermions get dynamical mass, they decouple for $E < m_F$ and coupling runs quickly and confines technicolor.

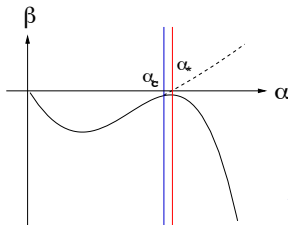


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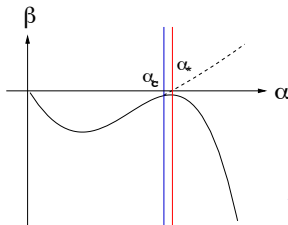


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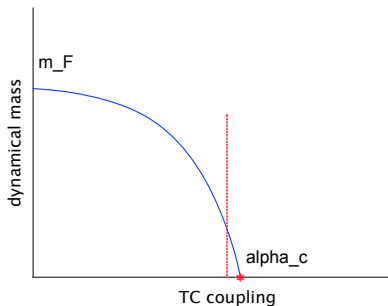
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- ▶ Large mass hierarchy due to quantum conformal phase transition at $\alpha = \alpha_c$: Miransky (or BKT) scaling near the phase transition

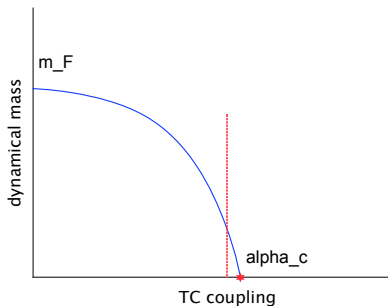


$$\gamma_{\bar{Q}Q} \simeq 1 + \sqrt{\frac{\alpha}{\alpha_c} - 1} \approx 1$$

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Very light techni-dilaton

- ▶ If $\alpha_* \approx \alpha_c$, theory exhibits walking behavior and is almost scale-invariant for a wide range of scale.
- ▶ Dilation current, $D^\mu = x_\nu \theta^{\mu\nu}$, is conserved up to scale-anomaly:

$$\langle \partial_\mu D^\mu \rangle = -\frac{\beta(\alpha)}{\alpha^2} \langle \alpha G_{\mu\nu}^a G^{a\mu\nu} \rangle_{\alpha \approx \alpha_*}^{TC} \approx 0.$$

- ▶ At some scale $\mu_{cr} \ll \Lambda_{TC}$, the coupling walks to cross the critical coupling $\alpha(\mu_{cr}) = \alpha_c$ and the theory undergoes a chiral phase transition:

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$$\langle 0 | D^\mu | \sigma \rangle = i F_{TD} p^\mu e^{-ip \cdot x}$$

- ▶ By PCDC

$$\langle \partial_\mu D^\mu \rangle = F_{TD} m_{TD}^2 \langle \sigma \rangle = F_{TD}^2 m_{TD}^2.$$

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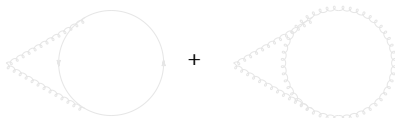
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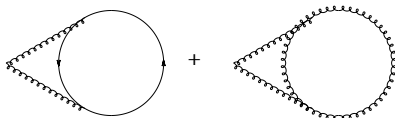
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- ▶ SD analysis by Hashimoto and Yamawaki shows as $\alpha \rightarrow \alpha_c$

$$\frac{m_F}{F_{TD}} \rightarrow 0.$$

- ▶ Since $\left(\frac{m_{TD}}{m_F}\right)^2 \sim \left(\frac{m_F}{F_{TD}}\right)^2$, TD is very light and decoupled:

$$m_{TD} \ll m_F (\approx 1 \text{ TeV}) \ll F_{TD}.$$

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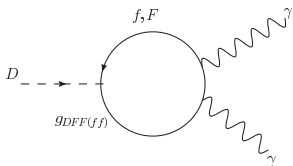
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Dark matter TD

► Decay of very light TD



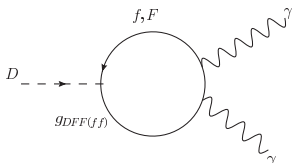
$$\Gamma(\sigma \rightarrow \gamma\gamma) \simeq \frac{\alpha_{em}^2}{36\pi^3} \frac{m_{TD}^3}{F_{TD}^2} |C|^2$$

$$\tau_{TD} \simeq 10^{17} \text{ sec } (N_{TC} N_F) \left(\frac{16}{c}\right)^2 \left(\frac{10 \text{ keV}}{m_{TD}}\right)^5 \left(\frac{m_F}{10^3 \text{ GeV}}\right)^4$$

- To be a dark matter candidate, TD has to be long-lived and $m_{TD} < 10 \text{ keV}$.

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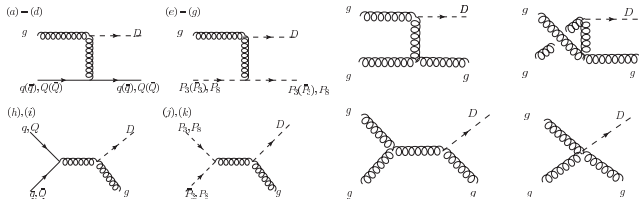
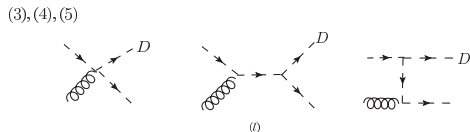
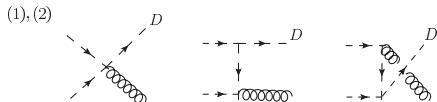
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Dark matter TD

► Thermal production of TD



Dark matter TD

- ▶ The Boltzmann equation for the TD number density n_{TD}

$$\frac{dn_{\text{TD}}}{dt} + 3Hn_{\text{TD}} = \sum_{i,j} \langle \sigma(i+j \rightarrow D + \dots) v \rangle n_i n_j$$

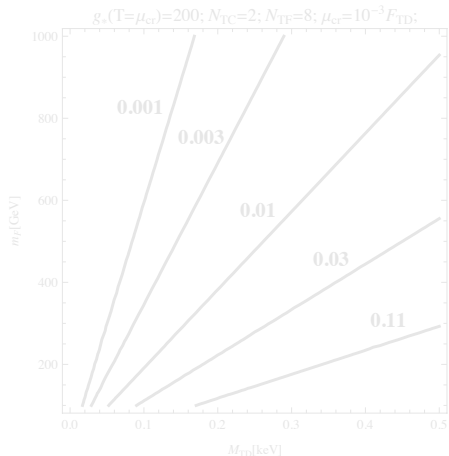
where H is the Hubble parameter $H(T) = (\frac{\pi^2}{30} g_* T^4 / 3M_{\text{P}}^2)^{1/2}$.

$$\Omega_{\text{TD}}^{\text{tp}} h^2 \simeq \left(\frac{\mu_{\text{cr}}}{10^8 \text{GeV}} \right) \left(\frac{m_{\text{TD}}}{\text{keV}} \right) \left(\frac{200}{g_*(\mu_{\text{cr}})} \right)^{3/2} \left(\frac{10^{11} \text{GeV}}{F_{\text{TD}}} \right)^2$$

$$\times \begin{cases} 4.4 \times 10^{-1} & \text{for } N_{\text{TC}} = 2 \\ 2.6 \times 10^{-2} & \text{for } N_{\text{TC}} = 3 \end{cases},$$

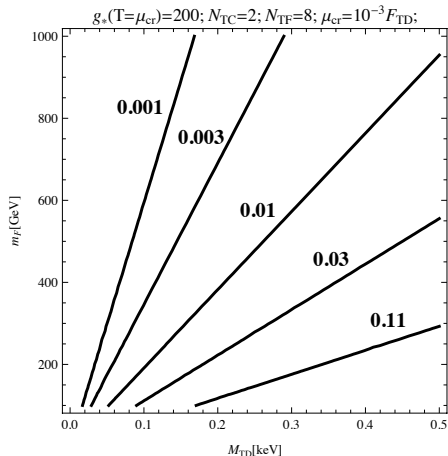
Dark matter TD

Contour plot of $\Omega_{TD}^{tp} h^2$
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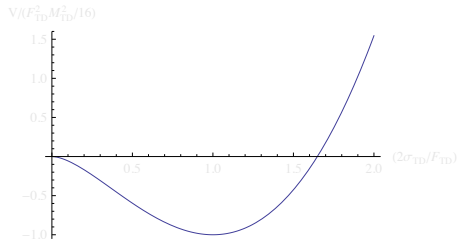


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- ▶ Non-thermal production of TD due to mis-alignment. TD potential is determined by scale anomaly (Schechter '80):

$$\frac{V(\sigma_D)}{F_{\text{TD}}^2 m_{\text{TD}}^2} \simeq \left(\frac{\sigma_D}{2F_{\text{TD}}} \right)^2 \left[\log \left(\frac{2\sigma_D}{F_{\text{TD}}} \right)^2 - 1 \right]$$

$$\theta \equiv \frac{\sigma_D}{\langle \sigma_D \rangle} - 1$$



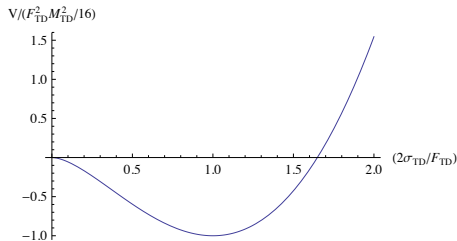
$$\Omega_{\text{TD}}^{\text{ntp}} h^2 \simeq 11 \times \left(\frac{\theta_{\text{os}}}{0.1} \right)^2 \left(\frac{200}{g_*(T_{\text{os}})} \right) \left(\frac{m_F}{10^3 \text{GeV}} \right)^4 \left(\frac{10^5 \text{GeV}}{T_{\text{os}}} \right)^3$$

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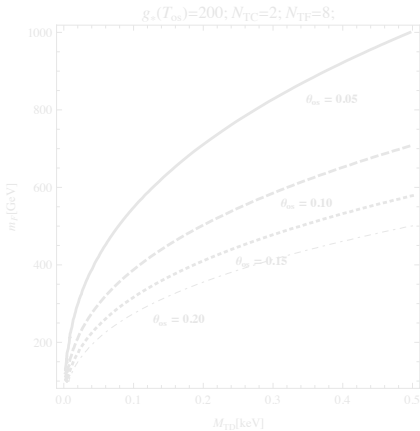
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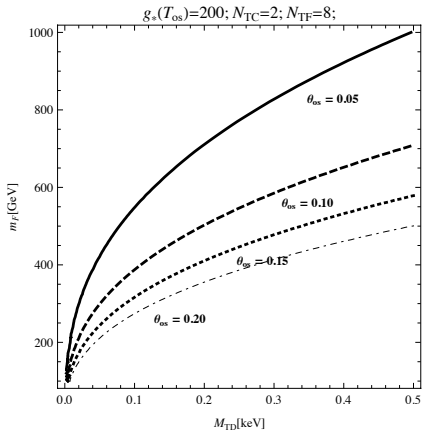
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- ▶ WTC predicts very light technidilaton (TD) due to spontaneously broken (approximate) scale symmetry.
- ▶ If the critical coupling, α_c , for chiral symmetry breaking is very close to the (quasi) Banks-Zaks IR fixed point of WTC, large hierarchy is dynamically generated:

$$m_F \approx \Lambda_{TC} e^{-\frac{\pi\alpha_c}{\sqrt{\alpha-\alpha_c}}}.$$

- ▶ By PCDC we have $m_{TD} \ll m_F \ll F_{TD}$.
- ▶ TD can be a good candidate for dark matter.
- ▶ Cosmological and astrophysical constraints require

$$0.01 \text{ eV} \lesssim m_{TD} \lesssim 500 \text{ eV}$$

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