

Quantum Field Theory: Example Sheet 1

Dr David Tong, October 2007

1. A string of length a , mass per unit length σ and under tension T is fixed at each end. The Lagrangian governing the time evolution of the transverse displacement $y(x, t)$ is

$$L = \int_0^a dx \left[\frac{\sigma}{2} \left(\frac{\partial y}{\partial t} \right)^2 - \frac{T}{2} \left(\frac{\partial y}{\partial x} \right)^2 \right] \quad (1)$$

where x identifies position along the string from one end point. By expressing the displacement as a sine series Fourier expansion in the form

$$y(x, t) = \sqrt{\frac{2}{a}} \sum_{n=1}^{\infty} \sin\left(\frac{n\pi x}{a}\right) q_n(t) \quad (2)$$

show that the Lagrangian becomes

$$L = \sum_{n=1}^{\infty} \left[\frac{\sigma}{2} \dot{q}_n^2 - \frac{T}{2} \left(\frac{n\pi}{a} \right)^2 q_n^2 \right]. \quad (3)$$

Derive the equations of motion. Hence show that the string is equivalent to an infinite set of decoupled harmonic oscillators with frequencies

$$\omega_n = \sqrt{\frac{T}{\sigma}} \left(\frac{n\pi}{a} \right). \quad (4)$$

2. Show directly that if $\phi(x)$ satisfies the Klein-Gordon equation, then $\phi(\Lambda^{-1}x)$ also satisfies this equation for any Lorentz transformation Λ .

3. The motion of a complex field $\psi(x)$ is governed by the Lagrangian

$$\mathcal{L} = \partial_\mu \psi^* \partial^\mu \psi - m^2 \psi^* \psi - \frac{\lambda}{2} (\psi^* \psi)^2. \quad (5)$$

Write down the Euler-Lagrange field equations for this system. Verify that the Lagrangian is invariant under the infinitesimal transformation

$$\delta\psi = i\alpha\psi \quad , \quad \delta\psi^* = -i\alpha\psi^* \quad (6)$$

Derive the Noether current associated with this transformation and verify explicitly that it is conserved using the field equations satisfied by ψ .

4. Verify that the Lagrangian density

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi_a \partial^\mu \phi_a - \frac{1}{2} m^2 \phi_a \phi_a \quad (7)$$

for a triplet of real fields ϕ_a ($a = 1, 2, 3$) is invariant under the infinitesimal $SO(3)$ rotation by θ

$$\phi_a \rightarrow \phi_a + \theta \epsilon_{abc} n_b \phi_c \quad (8)$$

where n_a is a unit vector. Compute the Noether current j^μ . Deduce that the three quantities

$$Q_a = \int d^3x \epsilon_{abc} \dot{\phi}_b \phi_c \quad (9)$$

are all conserved and verify this directly using the field equations satisfied by ϕ_a .

5. A Lorentz transformation $x^\mu \rightarrow x'^\mu = \Lambda^\mu_\nu x^\nu$ is such that it preserves the Minkowski metric $\eta_{\mu\nu}$, meaning that $\eta_{\mu\nu} x^\mu x^\nu = \eta_{\mu\nu} x'^\mu x'^\nu$ for all x . Show that this implies that

$$\eta_{\mu\nu} = \eta_{\sigma\tau} \Lambda^\sigma_\mu \Lambda^\tau_\nu. \quad (10)$$

Use this result to show that an infinitesimal transformation of the form

$$\Lambda^\mu_\nu = \delta^\mu_\nu + \omega^\mu_\nu \quad (11)$$

is a Lorentz transformation when $\omega^{\mu\nu}$ is antisymmetric: i.e. $\omega^{\mu\nu} = -\omega^{\nu\mu}$.

Write down the matrix form for ω^μ_ν that corresponds to a rotation through an infinitesimal angle θ about the x^3 -axis. Do the same for a boost along the x^1 -axis by an infinitesimal velocity v .

6. Consider the infinitesimal form of the Lorentz transformation derived in the previous question: $x^\mu \rightarrow x^\mu + \omega^\mu_\nu x^\nu$. Show that a scalar field transforms as

$$\phi(x) \rightarrow \phi'(x) = \phi(x) - \omega^\mu_\nu x^\nu \partial_\mu \phi(x) \quad (12)$$

and hence show that the variation of the Lagrangian density is a total derivative

$$\delta \mathcal{L} = -\partial_\mu (\omega^\mu_\nu x^\nu \mathcal{L}) \quad (13)$$

Using Noether's theorem deduce the existence of the conserved current

$$j^\mu = -\omega^\rho{}_\nu [T^\mu{}_\rho x^\nu] \quad (14)$$

The three conserved charges arising from spatial rotational invariance define the *total angular momentum* of the field. Show that these charges are given by,

$$Q_i = \epsilon_{ijk} \int d^3x (x^j T^{0k} - x^k T^{0j}) \quad (15)$$

Derive the conserved charges arising from invariance under Lorentz boosts. Show that they imply

$$\frac{d}{dt} \int d^3x (x^i T^{00}) = \text{constant} \quad (16)$$

and interpret this equation.

7. Maxwell's Lagrangian for the electromagnetic field is

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \quad (17)$$

where $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$ and A_μ is the 4-vector potential. Show that \mathcal{L} is invariant under gauge transformations

$$A_\mu \rightarrow A_\mu + \partial_\mu \xi \quad (18)$$

where $\xi = \xi(x)$ is a scalar field with arbitrary (differentiable) dependence on x .

Use Noether's theorem, and the spacetime translational invariance of the action, to construct the energy-momentum tensor $T^{\mu\nu}$ for the electromagnetic field. Show that the resulting object is neither symmetric nor gauge invariant. Consider a new tensor given by

$$\Theta^{\mu\nu} = T^{\mu\nu} - F^{\rho\mu} \partial_\rho A^\nu \quad (19)$$

Show that this object also defines four conserved currents. Moreover, show that it is symmetric, gauge invariant and traceless.

Comment: $T^{\mu\nu}$ and $\Theta^{\mu\nu}$ are both equally good definitions of the energy-momentum tensor. However $\Theta^{\mu\nu}$ clearly has the nicer properties. Moreover, if you couple Maxwell's Lagrangian to general relativity then it is $\Theta^{\mu\nu}$ which appears in Einstein's equations.

8. The Lagrangian density for a massive vector field C_μ is given by

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \frac{1}{2}m^2C_\mu C^\mu \quad (20)$$

where $F_{\mu\nu} = \partial_\mu C_\nu - \partial_\nu C_\mu$. Derive the equations of motion and show that when $m \neq 0$ they imply

$$\partial_\mu C^\mu = 0 \quad (21)$$

Further show that C_0 can be eliminated completely in terms of the other fields by

$$\partial_i \partial^i C_0 + m^2 C_0 = \partial^i \dot{C}_i \quad (22)$$

Construct the canonical momenta Π_i conjugate to C_i , $i = 1, 2, 3$ and show that the canonical momentum conjugate to C_0 is vanishing. Construct the Hamiltonian density \mathcal{H} in terms of C_0 , C_i and Π_i . (Note: Do not be concerned that the canonical momentum for C_0 is vanishing. C_0 is non-dynamical — it is determined entirely in terms of the other fields using equation (22)).

9. A class of interesting theories are invariant under the scaling of all lengths by

$$x^\mu \rightarrow (x')^\mu = \lambda x^\mu \quad \text{and} \quad \phi(x) \rightarrow \phi'(x) = \lambda^{-D} \phi(\lambda^{-1}x) \quad (23)$$

Here D is called the *scaling dimension* of the field. Consider the action for a real scalar field given by

$$S = \int d^4x \left[\frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2 - g \phi^p \right] \quad (24)$$

Find the scaling dimension D such that the derivative terms remain invariant. For what values of m and p is the scaling (23) a symmetry of the theory. How do these conclusions change for a scalar field living in an $(n+1)$ -dimensional spacetime instead of a $3+1$ -dimensional spacetime?

In $3+1$ dimensions, use Noether's theorem to construct the conserved current D^μ associated to scaling invariance.