

# Supergravity 1. Easter 2008

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1. In linearised supergravity the fields transform under a global supersymmetry transformation

$$h_{\mu\nu} \rightarrow h_{\mu\nu} - \frac{i}{2} \bar{\epsilon} (\gamma_\mu \psi_\nu + \gamma_\nu \psi_\mu)$$

$$\psi_\mu \rightarrow \psi_\mu - i \sigma^{\nu\rho} \partial_\nu h_{\rho\mu} \epsilon$$

Show that

$$[\delta_{\epsilon_1}, \delta_{\epsilon_2}] h_{\mu\nu} = -i (\bar{\epsilon}_2 \gamma^\rho \epsilon_1 \partial_\mu h_{\rho\nu} + \bar{\epsilon}_2 \gamma^\rho \epsilon_1 \partial_\nu h_{\rho\mu} - 2 \bar{\epsilon}_2 \gamma^\rho \epsilon_1 \partial_\rho h_{\mu\nu})$$

2. Use the above supersymmetry transformations but let  $\epsilon(x)$ , to show that the variation of the action for linearised gravity is

$$\delta S_{\text{linearised}} = \int d^4x j^\mu \partial_\mu \epsilon$$

where

$$j^\mu = \frac{i}{2} \varepsilon^{\mu\nu\rho\sigma} \bar{\psi}_\rho \gamma_5 \gamma^\nu \sigma^{\lambda\tau} \partial_\lambda h_{\tau\sigma}$$

3. Explain why the Rarita-Schwinger equation describes a spin 3/2 field. Compute the number of (off-shell) degrees of freedom for the massless gravitino and compare this to the corresponding number for the massless graviton. Show that this leads to the inclusion of auxiliary fields.

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4. Compute the number of on-shell degrees of freedom for a massless gravitino and a massless graviton in a D-dimensional spacetime, where D is even. When D=10 would you expect the supergravity multiplet to contain other dynamical fields?
5. The Wess-Zumino model is

$$\mathcal{L} = \partial_\mu \phi^* \partial^\mu \phi + \frac{i}{2} \bar{\psi} \gamma^\mu \partial_\mu \psi$$

where  $\psi$  is a Majorana spinor. The supersymmetry transformations are

$$\begin{aligned} \delta A &= \bar{\epsilon} \psi \\ \delta B &= i \bar{\epsilon} \gamma_5 \psi \\ \delta \psi &= -i \gamma^\mu \partial_\mu (A + i \gamma_5 B) \epsilon \end{aligned}$$

where

$$\phi = \sqrt{\frac{1}{2}} (A + iB)$$

when we let  $\epsilon$  be space-time dependent, i.e.  $\epsilon(x)$ , show that

$$\delta \mathcal{L} = \partial_\mu \bar{\epsilon} j^\mu$$

where

$$j^\mu = \not{\partial} (A - i \gamma_5 B) \gamma^\mu \psi$$

6. (\*) Show that if we add

$$\mathcal{L}_1 = \frac{-\kappa}{2} \bar{\psi}_\mu j^\mu$$

to the Lagrangian of question 5, then

$$\begin{aligned} \delta \mathcal{L}_1 &= -\partial_\mu \bar{\epsilon} j^\mu + i \kappa \bar{\psi}^\mu \gamma^\nu \epsilon T_{\mu\nu} \\ &\quad + i \frac{\kappa}{2} \epsilon^{\mu\nu\rho\sigma} \bar{\psi}_\mu \gamma_\nu \partial_\rho \epsilon A \overset{\leftrightarrow}{\partial}_\sigma B \\ &\quad + i \frac{\kappa}{2} \epsilon^{\mu\nu\rho\sigma} \partial_\rho \bar{\psi}_\mu \gamma_\nu \epsilon A \overset{\leftrightarrow}{\partial}_\sigma B \end{aligned}$$

where  $T_{\mu\nu}$  is the energy-momentum tensor for  $\mathcal{L}$  and

$$A \overset{\leftrightarrow}{\partial}_\sigma B = A \partial_\sigma B - B \partial_\sigma A$$

and we have used the leading order transformation

$$\psi_\mu \rightarrow \psi_\mu + \frac{2}{\kappa} \partial_\mu \epsilon$$