

The Reynolds Number

Consider a fluid of density ρ in which there is a characteristic speed U and a length scale L . Then it is natural to scale velocities by U and time by L/U .

Acceleration term of Navier-Stokes equation:

$$\rho \frac{\partial u}{\partial t} \sim \frac{\rho U}{L/U} \sim \frac{\rho U^2}{L}$$

Note: this is nonlinear in U

Viscous dissipation term:

$$\eta \frac{\partial^2 u}{\partial y^2} \sim \frac{\eta U}{L^2}$$



$$\frac{\text{inertial}}{\text{viscous}} = \frac{UL}{\nu} \equiv Re$$

$$\nu = \eta / \rho$$

Kinematic viscosity

For a bacterium in water, with $U \sim 10 \mu\text{m/s}$, $L \sim 1 \mu\text{m}$, and $\nu \sim 0.01 \text{ cm}^2/\text{s}$, $Re \sim 10^{-5}!!$

No Coasting at Low Reynolds Number

Ignoring any detailed fluid mechanics, we might imagine the equation of motion of a bacterium that has just switched off its flagellar motion to be of the form:

$$\frac{4}{3}\pi R^3 \rho \frac{d^2 x}{dt^2} = -6\pi\eta R \frac{dx}{dt}$$

Hence we deduce there is a characteristic time and distance

$$\tau \sim \frac{2}{9} \frac{R^2}{\eta/\rho}$$

$$\ell = v_0 \tau \sim \frac{2}{9} \frac{R^2 v_0}{\nu}$$

For a bacterium in water, with $v_0 \sim 10^{-3}$ cm/s, $R \sim 10^{-4}$ cm, and $\nu \sim 0.01$ cm²/s, $\tau \sim 10^{-7}$ s and $\ell \sim 10^{-10}$ cm!!

Typical forces and voltages

Let us estimate the typical force encountered by a swimming bacterium. Suppose it is in a fluid of viscosity η (0.01 for water), has a radius a on the order of a micron and swims at a speed v on the order of 10 $\mu\text{m/s}$. Using the Stokes drag law for a sphere the force on it is

$$F = 6\pi\eta av \simeq 2 \times 10^{-8} \text{dyne} \simeq 2 \times 10^{-13} \text{N} = 0.2 \text{ pN}$$

Forces on the cellular scale are on the order of pN. A very useful way to think about this is to convert thermal energy into pN nm

$$k_B T = 4 \times 10^{-14} \text{erg} \quad [k_B = 1.38 \times 10^{-16} \text{erg/K}]$$

$$k_B T = 4 \times 10^{-21} \text{J} = 4 \times 10^{-12} \text{N} \times 10^{-9} \text{m} = 4 \text{ pN} \cdot \text{nm}$$

With a very similar approach we can also estimate the typical voltages as

$$\frac{k_B T}{e} = \frac{4 \times 10^{-14} \text{erg}}{5 \times 10^{-10} \text{esu}} \simeq 25 \text{mV}$$

Advection & Diffusion

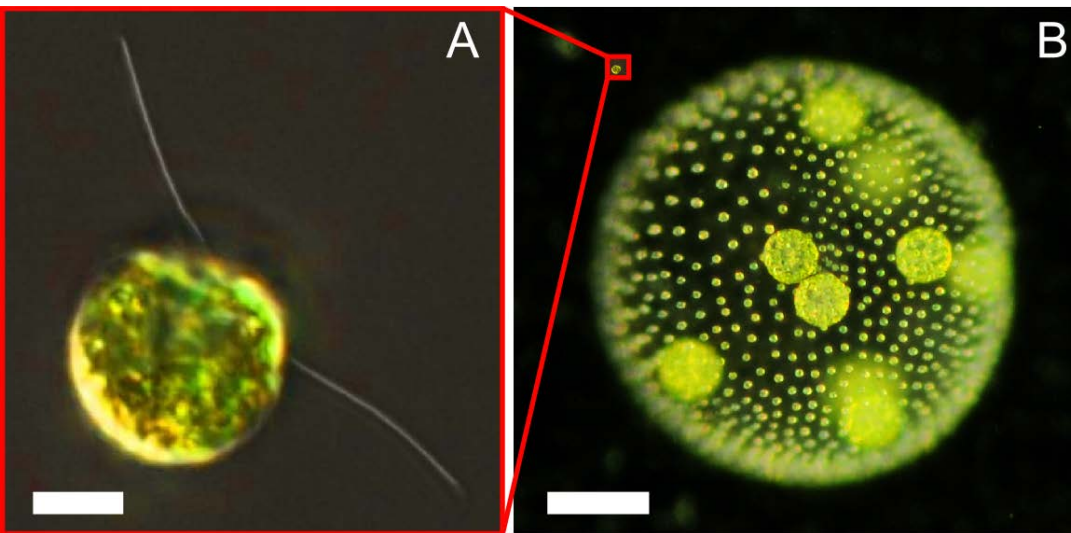
If a fluid has a typical velocity $\mathbf{U}$, varying on a length scale $\mathbf{L}$, with a molecular species of diffusion constant $\mathbf{D}$. Then there are two times:

We define the Péclet number as the ratio:

$$Pe = \frac{t_{diffusion}}{t_{advection}} = \frac{UL}{D}$$

This is like the Reynolds number comparing inertia to viscous dissipation:

$$t_{advection} = \frac{L}{U}$$
$$t_{diffusion} = \frac{L^2}{D}$$
$$Re = \frac{UL}{\nu}$$



If $U=10 \mu\text{m/s}$, $L=10 \mu\text{m}$,
 $Re \sim 10^{-4}$, $Pe \sim 10^{-1}$

*At the scale of an individual cell,
diffusion dominates advection.*

*The opposite holds for
multicellularity...*

Diffusion and the Stokes-Einstein Relation

If molecules have a diffusion constant D , concentration c , and are advected with speed u , then the flux is:

$$J = -D \frac{dc}{dx} + uc$$

In the low-Re regime we expect a force balance of the form $\zeta u = \text{force} = -d\phi/dx$, where ϕ is a suitable potential energy.

At equilibrium, we must have $J = 0$, so $0 = -D \frac{dc}{dx} - \frac{1}{\zeta} c \frac{d\phi}{dx}$, or

$$c \sim \exp(-\phi/D\zeta)$$

If equilibrium statistical mechanics holds then we must conclude that

$$D\zeta = k_B T \quad \text{or} \quad D = \frac{k_B T}{\zeta}$$

If ζ is the Stokes drag coefficient for a molecule of radius $2 \text{ \AA}$ we obtain

$$D \sim \frac{4 \times 10^{-14}}{20 \cdot 0.01 \cdot 2 \times 10^{-8}} \sim 10^{-5} \text{ cm}^2/\text{s}$$

Diffusional Time Scales

From the diffusion equation

$$\frac{\partial c}{\partial t} = D \nabla^2 c$$

we see by dimensional analysis the scaling

$$Dt \sim \ell^2 \quad \text{or} \quad t \sim \frac{\ell^2}{D}$$

On the scale of a bacterium ($\ell \sim 10^{-4}$ cm), $t \sim 10^{-3}$ s, but
on the scale of a plant ($\ell \sim 10$ cm), $t \sim 10^7$ s, or about 3-4 months!)

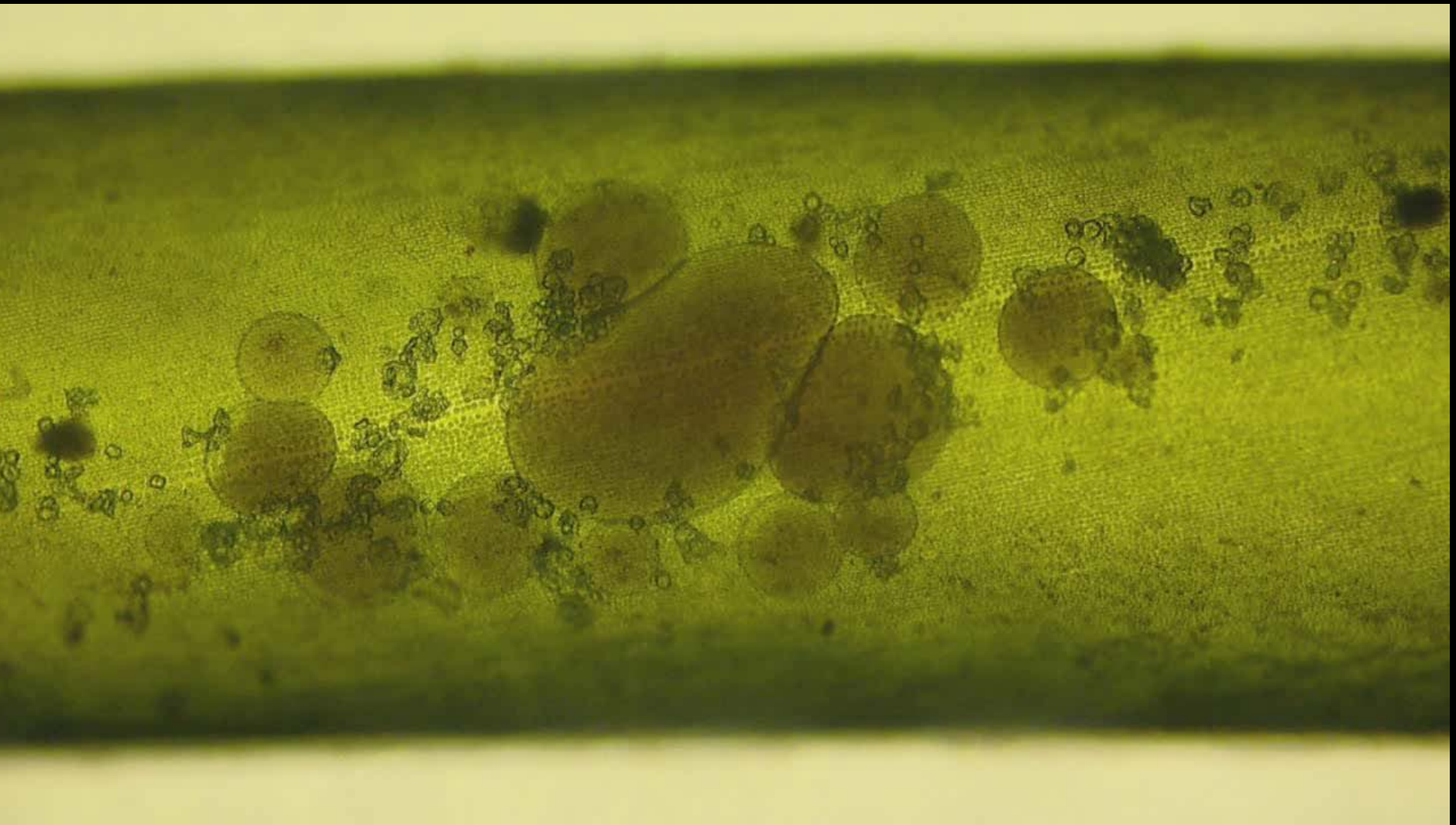
Something must take over as a transport mechanism beyond
Several hundred microns for life to function.

See J.B.S. Haldane, “On Being the Right Size”

The Aquatic Plant *Chara corallina*



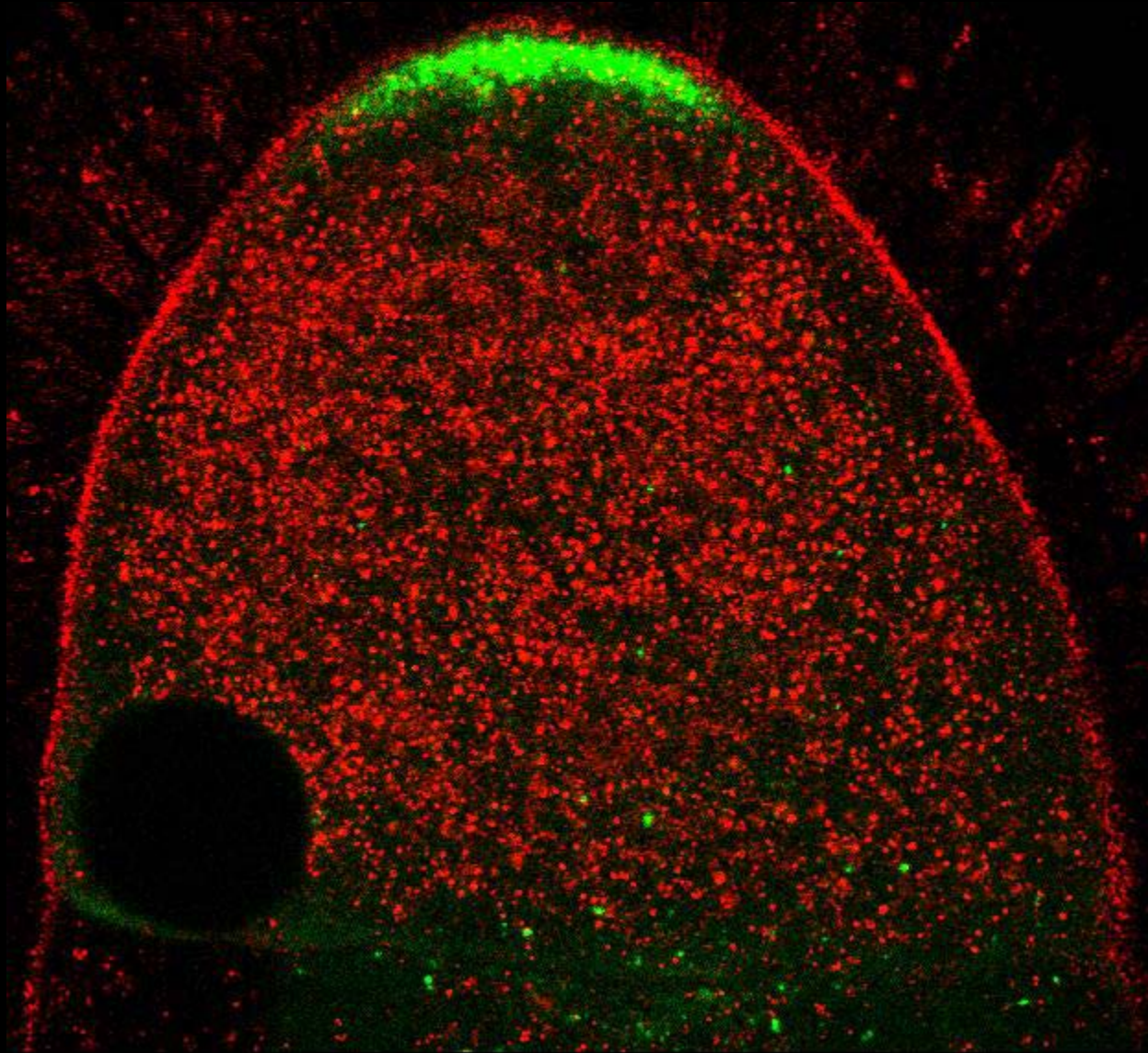
Cytoplasmic Streaming (the Movie)



Goldstein, Tuval, van de Meent, *PNAS* 105, 3663 (2008)

Streaming in *Drosophila* oogenesis

100 μm



Kinesin-based
movement of
cargo along
microtubules

speed ~ 30 nm/s

Viscosity very high

Ganguly, Williams, Palacios, Goldstein, *PNAS* **109**, 15109 (2012)

MKS (SI-units) $\Leftrightarrow$ CGS

- See pages 107-109 in “Van der Waals Forces” by Adrian Parsegian for full overview of conversion

- One example: Coulomb forces
in SI units

$$\text{Force} = \frac{q_1 q_2}{4\pi \epsilon_0 r^2} \text{ N, } q_1, q_2 \text{ in coulombs C, } r \text{ in meters,}$$

$$\epsilon_0 = 8.85 \times 10^{-12} \text{ C}^2 \text{ N}^{-1} \text{ m}^{-2}$$

or

$$(1/4\pi \epsilon_0) = 8.992 \times 10^9 \text{ N m}^2/\text{C}^2;$$

CGS units

$$\text{Force} = \frac{q_1 q_2}{r^2} \text{ dyn, } q_1, q_2 \text{ in statcoulombs, } r \text{ in centimeters}$$

- Second example: Poisson equation

$$\nabla \cdot (\epsilon \mathbf{E}) = \rho_{\text{free}}/\epsilon_0,$$

$$\nabla \cdot (\epsilon \mathbf{E}) = 4\pi \rho_{\text{free}}.$$

- Other conversion please see text book by A. Parsegian

Microscopics

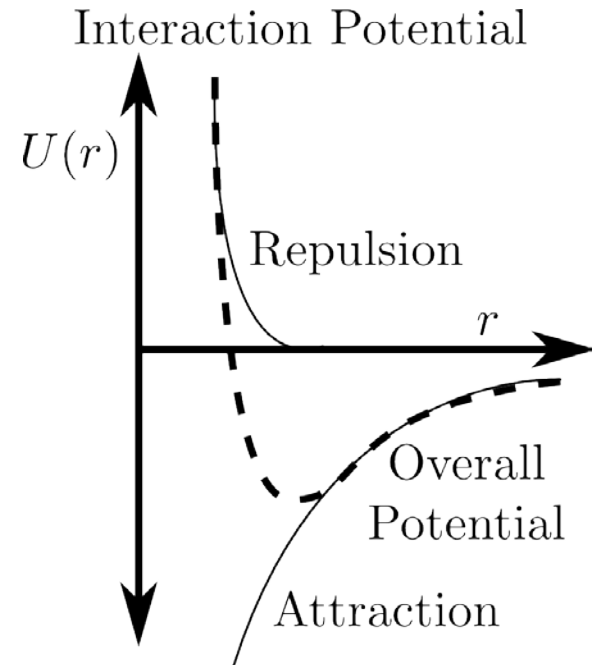
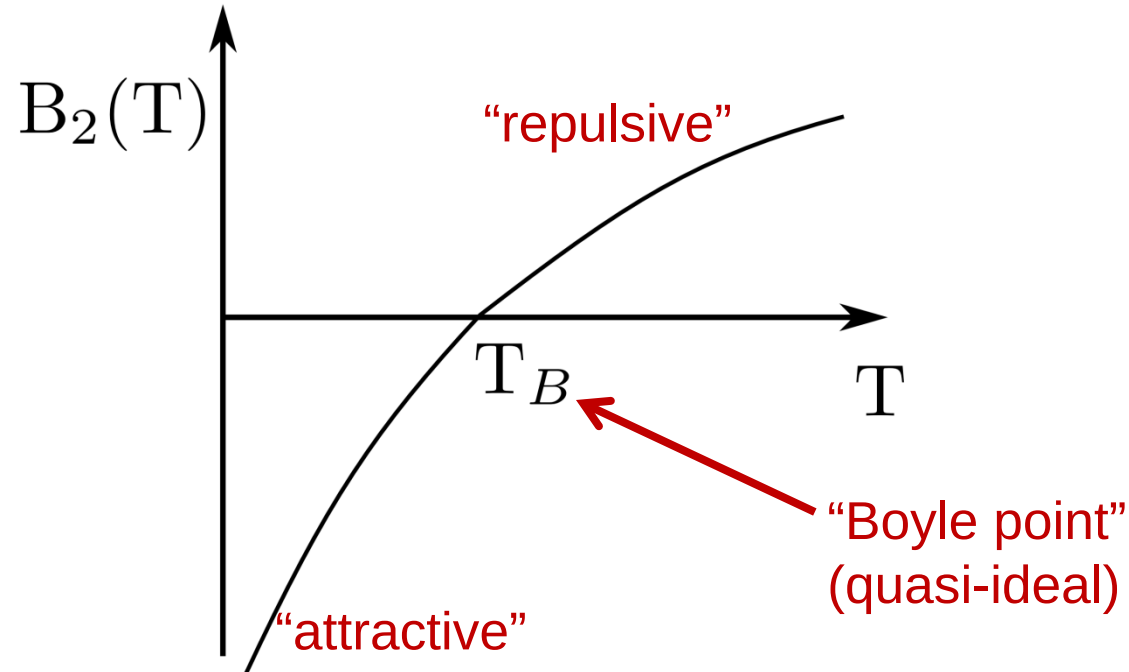
The ideal gas law,

$$PV = nRT = Nk_B T \qquad \frac{P}{k_B T} = \rho$$

is only an approximation. At low densities we have the "virial expansion",

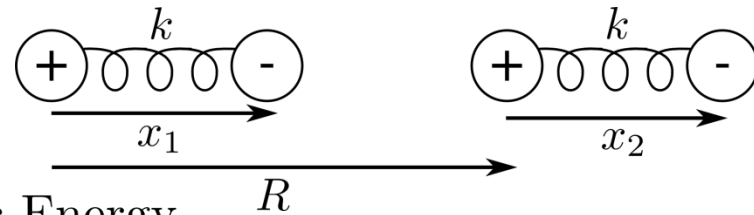
$$\frac{P}{k_B T} = \rho + B_2(T)\rho^2 + B_3(T)\rho^3 + \dots$$

2-body 3-body interactions



Van der Waals Interactions

The Hamiltonian for the system is:



$\mathcal{H} = \mathcal{H}_0 + \mathcal{H}_1 = \text{Spring Energy} + \text{Coulombic Energy}$

$$\mathcal{H}_0 = \frac{p_1^2}{2m} + \frac{1}{2}m\omega_0^2 x_1^2 + \frac{p_2^2}{2m} + \frac{1}{2}m\omega_0^2 x_2^2$$

$$\mathcal{H}_1 = e^2 \left[\frac{1}{R} + \frac{1}{R - x_1 + x_2} - \frac{1}{R - x_1} - \frac{1}{R + x_2} \right] \sim -\frac{2e^2 x_1 x_2}{R^3}, |x_1|, |x_2| \ll R$$

To simplify the situation, a coordinate change is made:

$$x_{\pm} = \frac{x_1 \pm x_2}{\sqrt{2}} \quad x_1 = \frac{x_+ + x_-}{\sqrt{2}} \quad x_2 = \frac{x_+ - x_-}{\sqrt{2}}$$

$$\mathcal{H} = \frac{p_1^2}{2m} + \frac{1}{2} \left(m\omega_0^2 - \frac{2e^2}{R^3} \right) x_+^2 + \frac{p_2^2}{2m} + \frac{1}{2} \left(m\omega_0^2 + \frac{2e^2}{R^3} \right) x_-^2$$

The two terms in parenthesis are essentially adjusted frequencies:

$$\omega_+^2 = \omega_0^2 - \frac{2e^2}{mR^3} \quad \omega_-^2 = \omega_0^2 + \frac{2e^2}{mR^3}$$

Van der Waals Interactions - continued

$$\omega_{\pm} = \left(\omega_0^2 \mp \frac{2e^2}{mR^3} \right)^{1/2} \simeq \omega_0 \mp \frac{e^2}{m\omega_0 R^3} - \frac{1}{2} \frac{e^4}{m^2 \omega_0^3 R^6} + \dots$$

The energy is then:

$$U(r) = \frac{1}{2} \hbar \omega_+ + \frac{1}{2} \hbar \omega_- - 2 \cdot \frac{1}{2} \hbar \omega_0$$
$$\approx \frac{-\hbar e^4}{2m^2 \omega_0^2 R^6} = \boxed{-\frac{1}{2} \hbar \omega_0 \frac{(e^2/m\omega_0^2)^2}{R^6}}$$

We see that $e^2/m\omega_0^2$ is a characteristic volume. To see the meaning of this, consider an induced electric dipole:

$$\mathbf{d} = \alpha \mathbf{E}$$

The units on $\mathbf{d}$ are $Q \cdot L$, the units on $\mathbf{E}$ are Q/L^2 , so the units of α are L^3 (a volume).

Van der Waals Interactions - continued

Now consider the Hamiltonian in the presence of an external electric field:

$$\begin{aligned}\mathcal{H} &= \mathcal{H}_0 + eE_0x_1 + eE_0x_2 \\ &= \frac{1}{2}m\omega_0^2 \left(x_1^2 + \frac{2eE_0x_1}{m\omega_0^2} \pm \left(\frac{eE_0}{m\omega_0^2} \right)^2 \right) + (1 \leftrightarrow 2) \\ &= \frac{1}{2}m\omega_0^2 z_1^2 + \dots \\ z_{1,2} &= x_{1,2} + \frac{eE_0}{m\omega_0^2}\end{aligned}$$

This is nothing but a shifted pair of harmonic oscillators, and from the shift we deduce the polarizability The dipole moment is then:

$$\alpha = \frac{e^2}{m\omega_0^2}$$

which is the same as seen previously. The energy of attraction in either case is then:

$$U(r) = -\frac{1}{2} \frac{\hbar\omega_0\alpha^2}{r^6}$$

How Does This Fit into the Thermodynamics?

Key idea [van der Waals, Weeks-Chandler-Anderson (WCA)]:
partition the intermolecular potential into purely repulsive + purely attractive,
use “known” results for former, perturbation theory for latter.

Using a mean-field (averaging) argument,

contribution to the
internal energy

$$U_{\text{attr}} = \frac{1}{2} N \rho \int d^3 r u_{\text{attr}}(r)$$


avoid double-counting

taking the density
to be uniform

Define $a = -\frac{1}{2} \int d^3 r u_{\text{attr}}(r)$ then $U_{\text{attr}} = -a N \rho = -\frac{a N^2}{V}$

$$\text{and } p_{\text{attr}} = -\frac{\partial U_{\text{attr}}}{\partial V} = -a \rho^2$$

So, a better equation of state is $p = \rho k_B T - a \rho^2$

To incorporate excluded volume, subtract Nb from the available volume,
where $b = 8 \times$ particle volume,

$$(p + a \rho^2)(V - Nb) = N k_B T$$

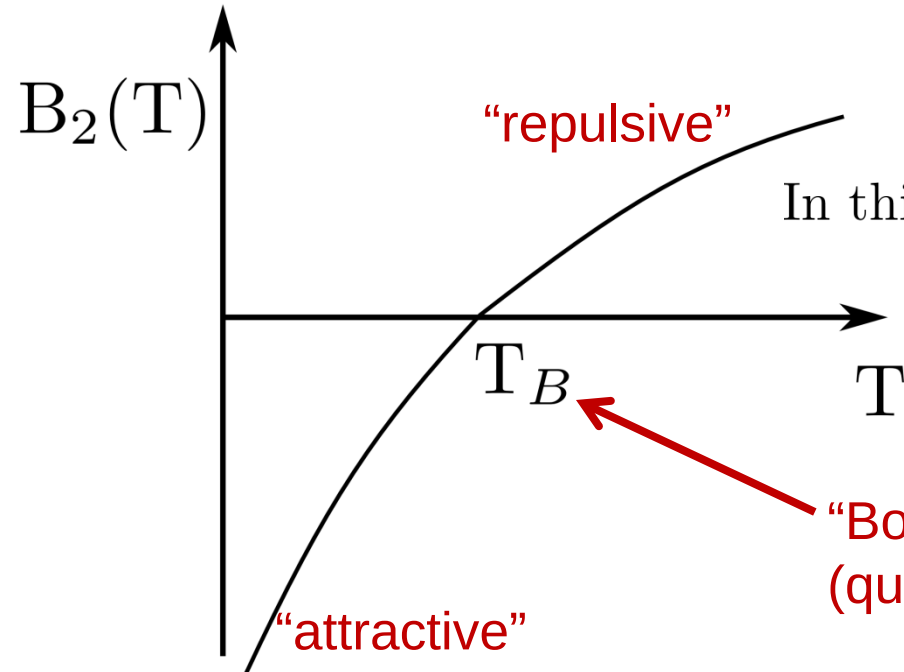
Putting it All Together

Finally, we have

$$p = \frac{\rho k_B T}{1 - b\rho} - a\rho^2 \quad \text{so} \quad \frac{p}{k_B T} \simeq \rho + \left(b - \frac{a}{k_B T}\right) \rho^2 + \dots$$

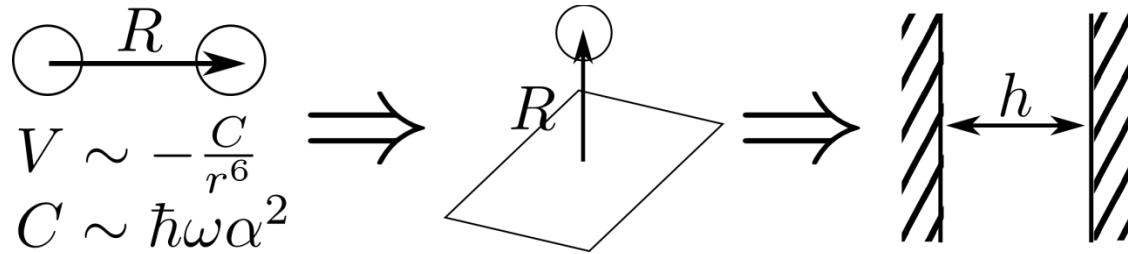
and we conclude that

$$B_2(T) = b - \frac{a}{k_B T}$$



$$T_B = \frac{a}{bk_B}$$

Applications of Dispersion Forces



Suppose the interaction between two molecules is

$$V_{11}(r) = -\frac{C}{r^6}$$

Then, the attraction between a neutral atom and a sheet is calculated:

$$\begin{aligned}
 V_{1S}(h) &= \int_h^\infty dz \int_0^{2\pi} d\phi \int_0^\infty r dr \rho V_{11}(\sqrt{z^2 + r^2}) \\
 &= - \int_h^\infty dz \int_0^{2\pi} d\phi \int_0^\infty r dr \rho \frac{C}{(z^2 + r^2)^3} \\
 &= -\frac{\pi C \rho}{6h^3} \quad (\text{atom-slab})
 \end{aligned}$$

Applications of Dispersion Forces - continued

Between two slabs of area A such that $A\rho dz$ atoms are in thickness dz :

$$V_{SS}(r) = A \int_h^\infty \rho V_{1S}(z) dz \Rightarrow \frac{V_{ss}}{A} = -\frac{A_H}{12\pi} \frac{1}{h^2}$$

This defines the Hamaker constant A_H

$$A_H = \pi^2 \rho_1 \rho_2 C_{12} \sim \pi^2 \hbar \omega (\alpha \rho)^2$$

Typically, the Hamaker constant is $\sim 5 \times 10^{-20} \text{ J} \sim 5 \times 10^{-13} \text{ erg}$,
About an order of magnitude larger than thermal energy

$$A_H \sim \pi^2 \hbar \omega_0 (\alpha \rho)^2 \sim 10 \cdot \text{eV} \cdot (0.1)^2 \sim 0.1 \text{ eV}$$