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# <sup>1</sup> Wave–vortex interactions and effective mean forces: three basic problems

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Three examples of wave-vortex interaction are studied, in analytically-tractable weak refraction regimes 5 with attention to the mean recoil forces, local and remote, that are associated with refractive changes in 6 wave pseudomomentum fluxes. Wave-induced mean forces of this kind can be persistent, with cumulative 7 8 effects, even in the absence of wave dissipation. In each example, a single wavetrain propagates past a single vortex. In the first two examples, in a two-dimensional, non-rotating acoustic or shallow-water setting, the 9 focus is on whether or not the wavetrain overlaps the vortex core. In the overlapping case, the recoil has 10 a local contribution given by the Craik-Leibovich force on the vortex core, the vector product of Stokes 11 drift and mean vorticity. (For a quantum vortex this contribution is called the Iordanskii force arising from 12 13 the Aharonov-Bohm effect on a phonon current.) However, in all except one special limiting case there are additional 'remote' contributions, mediated by Stokes-drift-induced return flows that can intersect the 14 vortex core well away from locations where the waves are refracted. 15

The third example is a non-overlapping, remote-recoil-only example in a rapidly rotating frame, in which the waves are deep-water gravity waves and the mean flow obeys shallow-water quasigeostrophic dynamics. Contrary to what might at first be thought, the Ursell 'anti-Stokes flow' induced by the rotation – an Eulerian-mean flow tending to cancel the Stokes drift – fails to suppress remote recoil. There are nontrivial open questions about extending these results to regimes of stronger refraction, especially regarding the scope of the 'pseudomomentum rule' for the wave-induced recoil forces.

Keywords: waves; momentum; pseudomomentum; wave-vortex interactions; mean recoil forces; missing
 forces in gravity-wave parametrizations; Iordanskii force; Aharonov-Bohm effect

# 24 1. Introduction

The following is a shortened, mainly descriptive version of a longer paper (McIntyre 2019, 25 hereafter M19), to which the reader is referred for full technical details. M19 explores ana-26 lytically tractable, precisely soluble versions of the wave–vortex interaction problems to be 27 discussed, with careful attention to asymptotic validity, and with cross-checks via indepen-28 dent analyses in complementary, but overlapping, asymptotic regimes. The present, shorter 29 paper tries to add value to M19 by providing a relatively concise summary of the main results 30 together with some additional discussion. The problems look simple at first sight but have 31 proved to be surprisingly tricky – conceptually as well as technically. They are fundamental, 32 moreover, to any attempt to complete our understanding of the  $O(a^2)$  wave-induced mean 33 forces arising from wave-induced momentum transport, where a is wave amplitude defined 34 such that  $a \ll 1$  validates linearization. 35

It hardly needs saying that mean forces of this kind are scientifically important. They 36 have long been recognized as playing a key role in, for instance, global-scale atmospheric 37 dynamics, as recalled in M19 and in greater detail in the reviews by Fritts (1984), Holton 38 et al. (1995), and Baldwin et al. (2001). See also Dritschel and McIntyre (2008, & refs.). 39 Examples include what used to be the enigma of the quasi-biennial oscillation or 13-monthly 40 reversal of the east-west winds in the equatorial lower stratosphere (e.g. Baldwin et al. 2001), 41 and the enigma of the cold summer mesopause with its noctilucent clouds (e.g. Fritts 1984). 42 (A global-scale circulation, gyroscopically pumped by wave-induced mean forces, turns the 43

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summer mesosphere into a giant refrigerator.) Some of the waves involved in these phenomena
are internal gravity waves whose scales are too small to be resolved in weather and climate
forecasting models, and whose mean effects are therefore routinely represented in the models
via so-called gravity-wave parametrization schemes (*e.g.* Garcia *et al.* 2017, & refs.).

The parametrization schemes and the associated theoretical literature have always, however, 48 neglected the nondissipative wave-induced mean recoil forces associated with the deflection 49 of waves by vortices, and other horizontal-refraction effects. Here, the word 'refraction' will 50 be used in its most general sense – which is the sense that is relevant here – to mean not 51 just the bending of rays but any distortion of the wave field by the vortex flow. Examples 52 include those illustrated in equations (2)-(5) below, as well as in studies like those of, for 53 instance, Sakov (1993), Sonin (1997), Coste et al. (1999) and Ford and Llewellyn Smith (1999, 54 hereafter FLS99). In various ways those examples include, but also go beyond, standard ray 55 theory (JWKB theory). The present work makes use of ray theory but also goes beyond it in 56 significant ways. 57

The nondissipative recoil forces in question are potentially important because they can 58 be persistent, in the same sense that the more familiar dissipative wave-induced forces are 59 persistent. They can act cumulatively over an arbitrary number of wave periods. And one of 60 the conceptually tricky questions about them is the question of *where* such forces are exerted. 61 Even the simplest problems, or thought-experiments, in which a single wavetrain is refracted 62 by a single vortex, illustrate what is involved. Consider the problem sketched schematically in 63 figure 1(a). A steady train of gravity waves or sound waves passes to one side of a vortex, in an 64 inviscid, two-dimensional, non-rotating shallow water or homentropic gas dynamical system. 65 The vortex flow has small Froude or Mach number 66

$$\epsilon = U/c_0 \ll 1, \tag{1}$$

and the wave refraction is correspondingly weak (and left invisible in the figure, but see section 2 below). Here U is a vortex flow speed and  $c_0$  an intrinsic wave speed. For definiteness,  $c_0$  will be taken as the wave speed at  $r = \infty$  and U as the flow speed at the edge of the vortex core,  $r = r_0$ , say, where  $r^2 = x^2 + y^2$  in the notation of the figure.

The question of where the mean recoil force is exerted is ambiguous. It can be asked and 71 answered in more than one way. Simplest and most useful is to ask the question in the way 72 that is relevant to gravity-wave parametrization. What force would be required if the waves 73 were removed, in order to have the same effect on the mean flow? For the problem sketched in 74 figure 1(a) the answer was found in an earlier study by Bühler and McIntyre (2003, hereafter 75 BM03). The answer may seem surprising at first sight. The force has to be exerted not where 76 the waves are refracted, within the wavetrain as it passes the vortex, but, rather, on the vortex 77 core. Because the core can be at an arbitrary distance from the wavetrain, BM03 called this 78 effective mean force a 'remote recoil'. 79

Of course there is no mystery here – no violation of Newton's Third Law – because a 80 fluid medium has a mean pressure field that can mediate actions and reactions continuously, 81 across substantial distances, just as in ordinary vortex-vortex interactions. The point may be 82 obvious, but is sometimes overlooked when problems like these are discussed from a particle-83 physics perspective. And it is perfectly reasonable to say, alternatively, that when the waves 84 are present the mean force is exerted where they are refracted. However, to make sense of the 85 resulting picture one would then have to solve for the  $O(a^2)$  mean pressure field and explicitly 86 describe how it transmits the force across the gap between the wavetrain and the vortex core. 87 BM03 also found that the mean force complies with what is now called the 'pseudo-88 momentum rule' (e.g. Bühler 2014, hereafter B14). Its validity is tacitly assumed in, for 89 instance, parts of the literature on gravity-wave parametrization, and on quantum vortices as 90 well (e.g. Sonin 1997). When valid, it avoids any consideration of the  $O(a^2)$  mean equations. 91 It says that the magnitude and direction of the mean force can be calculated from linearized 92



Figure 1. Panels (a) and (b) are schematics of wave–vortex interaction problems (i) and (ii) respectively. Waves of wavenumber  $\mathbf{k}$  are incident from the left and are weakly refracted by the vortex. The refraction effects are left invisible to emphasize their extreme weakness (but see section 2 below). The azimuthal angle  $\theta$  is defined unconventionally but in a way that will be convenient when discussing the Aharonov–Bohm effect, which turns out to be one of the significant wave refraction effects.

wave theory alone as if pseudomomentum were momentum, and as if the fluid medium were
 absent.

(Pseudomomentum, also called quasimomentum or wave momentum, or phonon momen-95 tum, is the  $O(a^2)$  linear-theoretic wave property whose nondissipative conservation depends, 96 through Noether's theorem, on translational invariance of the mean or background state on 97 which the waves propagate, as distinct from translational invariance of the entire physical sys-98 tem, background plus waves, which implies conservation of momentum (e.q. McIntyre 1981, 99 Peierls 1991, & refs.). In a linearized ray-theoretic description the pseudomomentum **p** per 100 unit mass is  $\mathcal{A}\mathbf{k}$ , where  $\mathbf{k}$  is the wavenumber vector and  $\mathcal{A}$  the wave-action, the intrinsic wave 101 energy divided by the intrinsic frequency, per unit mass. B14 gives more general expressions 102 for **p** outside the scope of linearized ray theory.) 103

Another way to state the pseudomomentum rule is to say that the magnitude and direction 104 of the recoil can be calculated as if the problem were one of particles such as photons hitting 105 an obstacle in a vacuum, *i.e.* ignoring the  $O(a^2)$  pressure field and treating the waves as 106 bullet-like, with momentum equal to their pseudomomentum. Such ideas are implicit in the 107 phraseology sometimes encountered in which the waves are described as exchanging 'their 108 momentum' with the mean state. A tendency to conflate momentum with pseudomomentum 109 can be found scattered throughout the physics literature under headings such as 'Abraham-110 Minkowski controversy' (e.g. Peierls 1991). 111

The problem sketched in figure 1(a) is the first of a set of three problems, or thoughtexperiments, considered here. It is shown in M19 that the pseudomomentum rule holds in all three of them, at least to leading order in  $\epsilon$ . The first two problems are in the two-dimensional, non-rotating setting and the third involves rotation:

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(i) As in figure 1(a). The vorticity  $\omega_0(r)$  is zero outside the vortex core.

(ii) As in figure 1(b). The wavetrain overlaps the vortex core. This brings in an additional
 refraction effect, familiar in the quantum literature as the so-called Aharonov–Bohm
 topological phase jump (figure 2 below).

(iii) As in figure 1(a) but in an inviscid, unstratified, incompressible, rapidly-rotating system of finite depth H with a free upper surface, under gravity g. The quasigeostrophic potential vorticity is uniform outside the vortex core. The waves are surface gravity waves with kH large enough to make  $\exp(-kH)$  negligible, where k is the magnitude of the wavenumber vector k. The mean-flow Rossby number is small, but the intrinsic wave frequency  $(gk)^{1/2} \gg f$ , the Coriolis parameter.

In all three problems it is assumed that  $a \ll \epsilon \ll 1$ , allowing linearized wave theory to be used

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to describe the weak wave refraction. In problem (iii) we take the mean-flow Rossby number to be of the same order as  $\epsilon$ . In some but not all cases, the wavenumber k is assumed large enough to permit the use of ray theory.

The plan of the paper is as follows. In section 2 and figure 2 we present and discuss a simple asymptotic solution for the O(a) wave field far from the vortex core, applicable to problems (i) and (ii) and describing the Aharonov–Bohm phase jump and other relevant wave-refraction effects. In a significant sense this solution encompasses ray theory but also goes beyond it.

In section 3 we introduce what turns out to be the simplest way to compute the mean recoil 134 forces and to understand their origin. It is to compute the complete nondissipative  $O(a^2)$ 135 mean flows associated with the wavetrains, sometimes called 'Bretherton flows'. Such a flow 136 consists of the Stokes drift within the wavetrain together with the return flow required by mass 137 conservation. For a narrow wavetrain the return flow takes place mostly outside it. Figure 3 138 shows an example. Section 4 presents the very simple mean-flow equations derived in M19. 139 They govern the Bretherton flows that are relevant to leading order in our three problems. 140 As noted in BM03 and in M19, this route to the results avoids any need to analyse the wave 141 refraction explicitly, a remarkable simplifying feature. To leading order in  $\epsilon$  it is enough to 142 compute Bretherton flows for the *unrefracted* wavetrains, *i.e.* correct to  $O(a^2\epsilon^0)$ . And that in 143 turn shows the leading-order results to be robust, in that they hold outside the ray-theoretic 144 and other regimes within which the wave refraction can be computed explicitly. 145

In section 5 we present the resulting formulae for the recoil forces in all three problems, 146 correct to leading order, in limiting cases for which the formulae become very simple. In 147 problem (ii) the simplest results are for wavetrains whose width W and length L are both 148 infinite. But immediately we encounter a surprise. The results depend strongly on the limiting 149 value of W/L. The limits  $W \to \infty$  and  $L \to \infty$  are noninterchangeable. Problem (iii) is 150 interesting in a different way. It exhibits remote recoil just as in problem (i), contrary to what 151 might be suggested by the effects of rapid rotation. Rotation produces a tendency for the 152 Stokes drift to be cancelled by the well-known 'anti-Stokes flow' (Ursell 1950). Nevertheless, 153 the cancellation is incomplete such that there is still a significant Bretherton flow giving rise to 154 a non-vanishing remote recoil, which moreover satisfies the pseudomomentum rule to leading 155 order in  $\epsilon$ . 156

Section 6 summarizes M19's independent and more lengthy, and indeed more delicate, 157 derivations of the same results from refraction calculations. Those derivations make use of 158 an appropriate 'impulse-pseudomomentum theorem', which justifies the pseudomomentum 159 rule independently, alongside direct calculations of the pseudomomentum fluxes in the re-160 fracted wave fields correct to  $O(a^2\epsilon^1)$ . Ray theory is used in some of these calculations, and 161 the non-ray-theoretic results of FLS99 in others. Section 7 offers brief concluding remarks 162 in which some challenges for future work are noted, particularly regarding what happens at 163 higher orders in  $\epsilon$ , for which the impulse-pseudomomentum theorem fails. In some but not all 164 circumstances the pseudomomentum rule still holds, but the precise circumstances remain to 165 be clarified. 166

## <sup>167</sup> 2. Wave refraction in problems (i) and (ii)

Before considering the  $O(a^2)$  mean flows, we note some key wave-refraction effects in problems (i) and (ii). As shown in M19, the linearized equations – see (2.1)–(2.3) of M19 – have the following far-field solution. For sufficiently large r, the velocity potential  $\phi'$  describing the waves has the asymptotic form  $\phi' = A \exp(i\Phi)$  where the O(a) amplitude envelope A is slowly-varying and where the phase  $\Phi$  is given by

$$\Phi = k_0(x - c_0 t) - \alpha \theta + \text{const.} + O(\epsilon^2 r_0^2 / r^2) .$$
<sup>(2)</sup>

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Figure 2. Wavecrests plotted from the far-field solution (2), with  $\alpha = 0.75$ . The unit of length is taken as  $k_0^{-1}$  so that the unrefracted wavelength is  $2\pi$ . The Aharonov–Bohm phase jump appears as a phase discontinuity on the positive x axis. In a full solution this discontinuity is smoothed out across a relatively narrow 'wake' region. The other relevant refraction effect is the very slight rotation of the wavecrests that can be seen, for instance, by careful inspection of the left-hand edge of the plot.

The incident wavenumber  $k_0$  and far-field phase speed  $c_0$  are constants, and  $\alpha$  is another constant, to be defined in (3). The azimuthal angle  $\theta$  is defined as in the figures. It ranges from  $-\pi$  to  $\pi$ .

In problem (ii), with the main focus on a wavetrain that is infinitely wide and infinitely long, we can take A to be a real constant. The constant  $\alpha$  in (2) is defined by

$$\alpha = \Gamma k_0 / 2\pi c_0 = U k_0 r_0 / c_0 = k_0 r_0 \epsilon \tag{3}$$

where  $\Gamma$  is the Kelvin circulation of the vortex, equal to  $\iint \omega_0 \, dx \, dy$ . The phase jump  $2\pi \alpha$  across 178 the positive x axis is the Aharonov–Bohm phase jump, a topological defect or dislocation of 179  $\Phi$ . In a full solution it is smoothed out across a relatively narrow 'wake' region surrounding 180 the positive x axis. The phase jump measures the effect of the vortex flow outside the core, 181  $\boldsymbol{u}_0(r) = Ur_0 r^{-1} \hat{\boldsymbol{\theta}} = \epsilon c_0 r_0 r^{-1} \hat{\boldsymbol{\theta}}$ , where  $\hat{\boldsymbol{\theta}}$  is the unit vector in the  $\boldsymbol{\theta}$  direction, in compressing the 182 wavetrain on one side while stretching it on the other, at positive and negative y respectively. 183 Figure 2 plots the far-field wavecrest shapes  $\Phi = \text{constant described by (2)}$  with the er-184 ror term neglected. We have taken  $\alpha = 0.75$ , fixing the phase jump at three quarters of a 185 wavelength to make it clearly visible. Also visible, less clearly, is another refraction effect that 186 nevertheless has comparable importance. Except on the y axis, the wavecrests are slightly 187 rotated away from the y direction, through angles  $O(\epsilon r_0 r^{-1})$ . The effect can be seen by care-188 ful inspection of the left-hand edge of the plot. The local wavenumber vector  $\boldsymbol{k} = \boldsymbol{\nabla} \Phi$  has a 189 refractive contribution  $-\alpha r^{-1}\hat{\theta} = -\epsilon k_0 r_0 r^{-1}\hat{\theta}$  directed against the vortex flow: 190

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where  $\hat{x}$  is the unit vector in the *x* direction. As in M19 we note that (2) and (4) are consistent with ray theory, but also go beyond it in the sense that phase changes over long distances of order *r* are represented accurately enough to describe the Aharonov–Bohm phase jump.

Despite the rotation of the wavecrests and of k, the absolute group velocity  $C^{abs}$  remains parallel to the x axis correct to  $O(\epsilon r_0 r^{-1})$ , as can be checked from its leading-order expression

$$\boldsymbol{C}^{\text{abs}} = = \frac{c_0 \boldsymbol{k}}{|\boldsymbol{k}|} + \boldsymbol{u}_0(r) + O(\epsilon^2 c_0 r_0^2 r^{-2})$$
(5)

<sup>196</sup> by taking the y components of (4) and of  $u_0(r) = \epsilon c_0 r_0 r^{-1} \hat{\theta}$ . Propagation due to the y <sup>197</sup> component of k cancels advection due to the y component of  $u_0$ . The cancellation follows <sup>198</sup> alternatively from the vanishing of the vorticity outside the vortex core, in virtue of the curl-<sup>199</sup> curvature formula of ray theory, B14 p. 86, which was first derived by Landau and Lifshitz <sup>200</sup> (1959) and generalized to dispersive waves by Dysthe (2001). Dysthe's result is made use of <sup>201</sup> in problem (iii).

The property of  $C^{abs}$  just noted means that (2) can also be applied to problem (i), with a *y*-dependent amplitude envelope A, as long as the wavetrain passes the vortex at a distance great enough for the expression (2) and ray theory to be asymptotically valid, as was assumed in BM03. Then A can be taken to depend on y alone in a way that restricts the wavetrain as sketched in figure 1(a). The width scale W of this envelope  $\gg k_0^{-1}$  for consistency with ray theory, yet small by comparison with the distance to the vortex core.

The phase function (2) is well known in the quantum literature. It applies not only to the vortex problem but also to the original Aharonov–Bohm problem, in which the waves represent nonrelativistic electrons going past a thin magnetic solenoid, as recalled in M19 Appendix A, with the magnetic vector potential in the role of the vortex flow  $\boldsymbol{u}_0(r)$ .

What is not apparent from (2) is the character of the wake region that smooths out the 212 Aharonov–Bohm phase jump. In the original Aharonov–Bohm problem the wake is symmetric 213 about the x axis, for arbitrary  $\alpha$ , and within it the discontinuous structure (2) is replaced by 214 a smooth Fresnel-diffractive structure with angular size tending toward zero like  $(k_0 x)^{-1/2}$  as 215  $x \to \infty$ . In the vortex problem, by contrast, the wake generally has small but non-vanishing 216 angular size  $O(\epsilon)$  and an asymmetry about the x axis of the same order, except in an extreme 217 long-wave limiting case with both  $\epsilon$  and  $k_0 r_0$  tending to zero. In that case the wake structure 218 tends toward a Fresnel-diffractive structure symmetric about the x axis, as shown in Sakov 219 (1993) and in FLS99. In cases of stronger refraction at finite  $\epsilon$ , the wave field becomes more 220 complicated and the wake asymmetry increased (Coste et al. 1999). In all cases, however, the 221 phase jump seen in figure 2 is smoothed out in some manner. 222

# 223 3. Bretherton flow and Kelvin impulse

Following the past literature including the pioneering work of Bretherton (1969), we use the term Bretherton flow to denote the entire  $O(a^2)$  wave-induced Lagrangian-mean flow. For instance in cases with relatively narrow wavetrains such as that of figure 1(a) the mean flow includes not only Stokes drifts but also any sideways return flows required by mass conservation. An example from BM03 is shown in figure 3. This represents schematically a version of problem (i) analysed in section 5.1 of BM03, in a particular formal limit, namely that of an infinitely long wavetrain slightly deflected by the vortex.

Within the wavetrain (which again is considered wide by comparison with  $k_0^{-1}$ , like a laser beam, even though narrow by comparison with the distance to the vortex core), the Stokes drift is toward the right. Therefore the return part of the Bretherton flow advects the vortex core toward the left. The core translates leftward at velocity  $u_{\rm tr}$  say. The resulting rate of change in the Kelvin impulse I of the vortex, eq. (7) below, is the same as if the waves

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Figure 3. Schematic of the Bretherton flow arising in a version of problem (i) studied in BM03. The  $O(a^2)$  mean flow within a narrow wavetrain, whose ray path is shown by the heavy curve, is dominated by the Stokes drift. A small portion of its mass flux,  $O(a^2\epsilon^1)$  in this case, leaks sideways as a consequence of wave refraction. To describe this situation the refraction problem must be considered correct to two orders in  $\epsilon$ , as was done in section 5.1 of BM03. Refraction effects enter at both orders, not only the  $O(\epsilon)$  effects illustrated in figure 2 but also an  $O(\epsilon^2)$  change in the direction of the absolute group velocity, exaggerated in this schematic.

were removed and a suitably tailored body force field F, pointing in the +y direction, were artificially applied to the vortex core. As already indicated, this is the effective mean recoil force in the sense considered here. It is exactly the sense required by – though, in fact, so far neglected in – gravity-wave parametrizations in weather and climate forecasting models.

The 'remoteness' of the recoil can now be seen to be related to the fact that the return flow extends well outside the wavetrain in cases like this. The Stokes drift does not directly contribute to  $u_{tr}$ , but only the return part of the Bretherton flow. In problem (ii), by contrast, the wavetrain overlaps the vortex core so that the local Stokes drift  $\overline{u}^{S}$  contributes to  $u_{tr}$ , as well as remote contributions from other parts of the wavetrain.

For our core with vorticity  $\omega_0(r)$  it is easy to verify that the effective force is just  $\mathbf{F} = -\omega_0 \hat{\boldsymbol{z}} \times \boldsymbol{u}_{tr}$  where the unit vector  $\hat{\boldsymbol{z}}$  points out of the paper.<sup>1</sup> Being transverse to the vortex motion, the resultant force  $\boldsymbol{R}$  has the character of a Magnus force, namely

$$\boldsymbol{R} = \iint \boldsymbol{F} \, \mathrm{d}x \mathrm{d}y = -\boldsymbol{\hat{z}} \times \boldsymbol{u}_{\mathrm{tr}} \iint \omega_0 \, \mathrm{d}x \mathrm{d}y = -\Gamma \boldsymbol{\hat{z}} \times \boldsymbol{u}_{\mathrm{tr}} \,. \tag{6}$$

We note that the Kelvin circulation  $\Gamma$  is an  $O(\epsilon)$  quantity and that, in the case of figure 3, BM03 found that  $\boldsymbol{u}_{tr}$  is  $O(a^2\epsilon^1)$  so that  $\boldsymbol{R}$  is  $O(a^2\epsilon^2)$ . It is readily shown (M19, eq. (3.7)) that  $d\boldsymbol{I}/dt = \boldsymbol{R}$  for our translating vortex core, with vorticity  $\omega_0(r')$  where  $r' = |\boldsymbol{x} - \boldsymbol{u}_{tr}t|$ . The two-dimensional Kelvin impulse  $\boldsymbol{I}$  is defined by

$$\boldsymbol{I} = \iint (y, -x) \,\omega_0 \,\mathrm{d}x \mathrm{d}y = \iint -\hat{\boldsymbol{z}} \times \boldsymbol{x} \,\omega_0 \,\mathrm{d}x \mathrm{d}y \tag{7}$$

(e.g. Batchelor 1967, equation (7.3.7)). In the case of figure 3, BM03 also found that  $u_{\rm tr}$  is just such that R satisfies the pseudomomentum rule.

The effect on the vortex, continually moving it parallel to the x axis, is persistent and cumulative, over an arbitrary number of wave periods. In that respect the wave-induced recoil

<sup>&</sup>lt;sup>1</sup>The curl of this two-dimensional force field  $\mathbf{F}$  is just that required to move the vortex core leftward through the fluid at velocity  $\mathbf{u}_{tr}$ , while the divergence of  $\mathbf{F}$  sets up the dipolar pressure field required to produce the corresponding changes outside the core, where the velocity field is irrotational. Thus defined,  $\mathbf{F}$  has the dimensions of acceleration, length/(time)<sup>2</sup>, *i.e.* force per unit mass, since it is a forcing term on the right-hand side of the standard momentum equation having  $\partial \mathbf{u}/\partial t$  on the left, whose curl is the standard vorticity equation. So for instance the resultant force on a two-dimensional vortex core of depth H is  $\rho H \iint \mathbf{F} dxdy$  where  $\rho$  is fluid density. The factor  $\rho H$  will be ignored in what follows. Strictly speaking, therefore, 'resultant force' and 'impulse' in the main text should be read as  $\rho^{-1}$  times resultant force and impulse per unit core depth.

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is like the wave-induced mean forces that arise from wave dissipation, even though in our three 256 problems there need not be any such dissipation. Even the wave sink need not be dissipative. 257 It can be a wavemaker whose amplitude and phase are contrived to give perfect absorption, 258 as for instance in the thought-experiments used by Léon Brillouin in his classic works on 259 radiation stress (e.g. Brillouin 1936, & refs.). 260

Two further points to note are first that the wave field can be taken as steady only as an 261 approximation, for small a and  $\epsilon$ , and second that the weakness,  $O(a^2)$ , of the return flow 262 and its strain-rate means that the vortex core is advected bodily without significant distortion 263 (Kida 1981) as indeed was already assumed below (6), simplifying the calculation of dI/dt264 from (7). 265

In problem (ii), as already said, there is an additional, local contribution to  $u_{\rm tr}$  and therefore 266 to R, from the Stokes drift  $\overline{u}^{S}$  of the wavetrain where it overlaps the vortex core. This local 267 contribution is just the Craik–Leibovich vortex force as usually defined,  $F_{\rm CL} = \overline{u}^{\rm S} \times \omega_0$ , 268 where  $\boldsymbol{\omega}_0 = \omega_0 \hat{\boldsymbol{z}}$ . The remote or return-flow contribution, from other parts of the wavetrain, 269 varies with W/L. So it is the remote and not the local contribution that gives rise to the 270 noninterchangeability of limits already mentioned. 271

#### Mean-flow equations at leading order 272 4.

From here on we restrict attention to leading-order,  $O(a^2\epsilon^1)$  recoil forces, thus excluding 273 further consideration of cases like that of figure 3 in which the recoil is  $O(a^2\epsilon^2)$  or smaller. 274 Then (6) can be used with  $u_{\rm tr}$  correct to  $O(a^2\epsilon^0)$  only, because of the factor  $\Gamma = O(\epsilon)$ . So as 275 said earlier we need only compute Bretherton flows for unrefracted wavetrains. 276

As shown in M19, at this order the Bretherton flows  $\overline{u}_{\rm B}^{\rm L}$  are nondivergent, with streamfunc-277 tion 278

$$\tilde{\psi}_{\rm B} = \tilde{\psi} - \tilde{\psi}_0 \tag{8}$$

say, where  $\tilde{\psi}_0 = \tilde{\psi}_0(r')$  is the streamfunction for the nondivergent velocity field  $u_0(r')$  of the vortex flow, and where the complete Lagrangian-mean flow  $\overline{u}^L$ , vortex flow  $u_0$  plus Bretherton 279 280 flow  $\overline{\boldsymbol{u}}_{\mathrm{B}}^{\mathrm{L}}$ , has x and y components 281

$$\overline{u}^{\mathrm{L}} = -\frac{\partial \tilde{\psi}}{\partial y} \quad \text{and} \quad \overline{v}^{\mathrm{L}} = \frac{\partial \tilde{\psi}}{\partial x}.$$
 (9)

M19 showed that correct to  $O(a^2\epsilon^0)$  the mean-flow equations can be written in the very simple 282 forms 283

$$\nabla_{\rm H}^2 \tilde{\psi}_{\rm B} = \hat{\boldsymbol{z}} \cdot \boldsymbol{\nabla} \times \boldsymbol{\mathsf{p}}$$
 in problems (i) and (ii) (10)

and 284

$$(\nabla_{\rm H}^2 - L_{\rm D}^{-2})\tilde{\psi}_{\rm B} = \hat{\boldsymbol{z}} \cdot \boldsymbol{\nabla} \times \langle \boldsymbol{\mathsf{p}} \rangle \qquad \text{in problem (iii)}, \qquad (11)$$

where  $\mathbf{p}$  is the wavetrain's pseudomomentum per unit mass, as before,  $\nabla_{\rm H}^2$  is the Laplacian in 285 the xy plane, and  $L_{\rm D}$  in problem (iii) is the Rossby deformation length-scale  $L_{\rm D} = f^{-1}(gH)^{1/2}$ . 286 The angle brackets denote vertical averaging, needed in problem (iii) because of the strong 287 dependence of **p** upon the vertical coordinate z, namely  $\mathbf{p} \propto \exp(2k_0 z)$ . The streamfunction 288  $\psi_{\rm B}$  need not be averaged vertically, in problem (iii), because at small Rossby number the 289 Taylor–Proudman effect makes it z-independent. 290

The simplicity of eqs. (10) and (11) comes from their close relation to Kelvin's circulation 291 theorem as expressed most succinctly by GLM (generalized Lagrangian-mean) theory; see for 292 instance section 10.2.7 of B14, and equations (2.9)-(2.11) of M19. 293

Because the wavetrains are unrefracted, they can be taken to have the simple sinusoidal 294 structure  $A \exp(i\Phi)$  with  $\Phi = k_0(x - c_0 t)$ , and constant or slowly-varying A. From this, and 295 from the irrotationality of the wavemotion in all three problems, we have  $\mathbf{p} = \overline{\mathbf{u}}^{S}$ . See for 296 instance B14, eqs. (10.15) and (10.17). The elliptic operators on the left of (10) and (11) show 297 why Bretherton flows extend well outside any relatively narrow wavetrain. 298



Figure 4. Schematic of Bretherton-flow streamlines in problem (i), as analysed in BM03 correct to lowest order  $O(a^2\epsilon^0)$  for the finite wavetrain whose ray path is shown by the heavy straight line. At this order the Stokes drift is nondivergent except within the wave source and sink regions. The waves propagate from a source on the left to a sink on the right.

### <sup>299</sup> 5. Bretherton flows and recoil forces at leading order

To take advantage of the simplifications just noted, with Bretherton flows computed from 300 unrefracted wavetrains, we need to consider a wavetrain of finite length in the case of problem 301 (i) as was done in BM03. In the formal limit of an infinitely long wavetrain, in that problem, 302 to leading order in  $\epsilon$ , the recoil and net pseudomomentum flux vanish because the rays remain 303 straight and parallel to the x axis thanks to the cancellation already noted in the absolute 304 group velocity (5), between the y components of the leading terms. (The bending of rays 305 indicated in figure 3 takes place at the next order in  $\epsilon$ , with the mean-flow equations becoming 306 less simple, though still elliptic, as shown in section 5.1 of BM03.) 307

So for problem (i) at leading order we consider the situation sketched in figure 4, with an unrefracted wavetrain of finite length marked by the heavy straight line, along with its surrounding return flow satisfying eq. (10). The heavy straight line corresponds to the wavetrain sketched in figure 1(a), whose width scale  $W \gg k_0^{-1}$  to permit the use of ray theory as in BM03. So once again the wavetrain is wide by comparison with  $k_0^{-1}$ , like a laser beam, even though narrow by comparison with the distance to the vortex core.

The wave source and sink are modelled as irrotational body-force fields, with  $k_0$  large enough 314 to allow the wave source and sink to be considered approximately localized near positions 315  $(x,y) = (\pm X,Y)$ , say, with  $X,Y \gg W$ . The Stokes drift is toward the right, straight along 316 the wavetrain where the right-hand side of (10) is nonzero, and the return flow with right-317 hand side zero is irrotational, emanating from the wave sink and returning through the wave 318 source. Its streamlines are mirror-symmetric about the wavetrain, because the wavetrain is 319 unrefracted at this order. As in BM03 and M19 the  $O(a^2\epsilon^0)$  flow advecting the vortex core at 320 (x, y) = (0, 0) can then be shown to be 321

$$\boldsymbol{u}_{\rm tr} = \overline{\boldsymbol{u}}_{\rm B}^{\rm L}(0,0) = \frac{S}{\pi} \frac{X}{X^2 + Y^2}(-\hat{\boldsymbol{x}}) ,$$
 (12)

322 where

$$S = \int \mathsf{p}_1(y) dy = \int \overline{u}^{\mathrm{S}}(y) dy \ . \tag{13}$$

The pseudomomentum per unit mass within the wavetrain has been written as  $\mathbf{p} = \mathbf{p}_1(y)\hat{x} = \overline{u}^{\mathrm{S}}(y)\hat{x}$ , and the integral is taken across the wavetrain. Applying the Magnus formula (6) we see that the corresponding recoil force is

$$\boldsymbol{R} = \frac{\Gamma S}{\pi} \frac{X}{X^2 + Y^2} (+ \hat{\boldsymbol{y}}) \tag{14}$$

<sup>326</sup> correct to  $O(a^2\epsilon^1)$ , where  $\hat{\boldsymbol{y}}$  is the unit vector in the *y* direction. As shown in BM03 and in <sup>327</sup> B14 there is a corresponding imbalance between the mean forces exerted by the wave source <sup>328</sup> and sink, as implied by Newton's third law. 10

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We note that (12) and (14) tend toward zero in the formal limit of an infinitely long wavetrain,  $X \to \infty$ . More precisely, since (13) implies the scaling  $S \sim W|\overline{\boldsymbol{u}}^{\mathrm{S}}|$ , we have

$$|\boldsymbol{u}_{tr}| \sim W|\overline{\boldsymbol{u}}^{S}|/X \text{ and } |\boldsymbol{R}| \sim \Gamma W|\overline{\boldsymbol{u}}^{S}|/X \text{ as } X \to \infty,$$
 (15)

going to zero like  $X^{-1}$ . The irrotational return part of the Bretherton flow becomes increasingly spread out in the y direction, diluting its effect at (x, y) = (0, 0). The vanishing of  $\mathbf{R}$  in this formal limit in problem (i) is consistent with the straightness of the rays and the vanishing of all refractive distortions as  $|x| \to \infty$ , at fixed y = Y in (2). In the limit the incoming and outgoing pseudomomentum fluxes become equal, and the wave source and sink exert equal and opposite mean forces.

The dilution effect summarized by (15) is key to understanding the noninterchangeability of limits in problem (ii). For instance, the same dilution effect occurs for any unrefracted wavetrain whose width W is given an arbitrary fixed value while its length  $L = 2X \rightarrow \infty$ , whether or not it overlaps the vortex core. When it does overlap, the local Stokes drift  $\overline{\boldsymbol{u}}^{S}$ contributes to  $\boldsymbol{u}_{tr}$  while the diluted return flow is still governed by (15), going to zero in the limit. It remains zero if the limit  $W \rightarrow \infty$  is taken subsequently. Therefore, for problem (ii) in the limit  $L \rightarrow \infty$  followed by  $W \rightarrow \infty$ , the formulae (12) and (14) are replaced by

$$\boldsymbol{u}_{\rm tr} = \overline{\boldsymbol{u}}^{\rm S}(0,0) = \boldsymbol{\mathsf{p}}_1(0)(+\boldsymbol{\hat{x}}) \tag{16}$$

$$\boldsymbol{R} = \Gamma \boldsymbol{\mathsf{p}}_1(0)(-\boldsymbol{\hat{y}}) . \tag{17}$$

Not only the magnitudes but also the signs have changed. Notice again that (17) is equal to the Craik–Leibovich vortex force  $\mathbf{F}_{CL} = \overline{\boldsymbol{u}}^S \times \boldsymbol{\omega}_0$  integrated over the vortex core, corresponding to what is called the Iordanskii force in the quantum vortex literature (*e.g.* Sonin 1997, Stone 2000), with  $\overline{\boldsymbol{u}}^S = \boldsymbol{p}$  corresponding to the phonon current per unit mass.

If we take the limits in the opposite order,  $W \to \infty$  followed by  $L \to \infty$ , it is easy to see that  $u_{tr}$  and R both go to zero. For an infinitely wide wavetrain of finite length, with  $\mathbf{p}_1$  uniform across the wavetrain, the dilution effect is banished to  $|y| = \infty$  so that the return flow at each finite |y| is just  $-\overline{u}^S$ . Thus  $u_{tr} = \overline{u}^S - \overline{u}^S = 0$ . For intermediate cases in which W/L has a finite limiting value, and in which the wavetrain is uniform and symmetric about the |x| axis, we obtain the intermediate values

$$\boldsymbol{R} = -\left\{1 - \frac{2}{\pi} \lim \arctan\left(\frac{W}{L}\right)\right\} \Gamma \boldsymbol{p}_1 \boldsymbol{\hat{y}} , \qquad (18)$$

as shown in M19, where lim denotes the limit  $W \to \infty$  and  $L \to \infty$  with W/L tending to a constant. In cases where the constant has a value of order unity, the local and remote contributions have comparable importance.

In problem (iii) there is no dilution effect as  $L \to \infty$ , because the Bretherton flow satisfies (11) and therefore, for a long wavetrain, decays sideways like  $\exp(-|y|/L_{\rm D})$  on the fixed lengthscale  $L_{\rm D}$ . In the formal limit  $L \to \infty$ , and with a narrow wavetrain,  $W \ll Y$  and  $W \ll L_{\rm D}$ , we have  $\overline{\boldsymbol{u}}_{\rm B}^{\rm L} = \overline{\boldsymbol{u}}_{\rm B}^{\rm L}(y) = (S/2L_{\rm D})\exp(-|y-Y|/L_{\rm D})(-\hat{\boldsymbol{x}})$  for |y-Y| > W, *i.e.* outside the wavetrain, so that, for a small vortex core  $r_0 \ll L_{\rm D}$ , the vortex-core advection velocity and recoil force are

$$\boldsymbol{u}_{\rm tr} = \overline{\boldsymbol{u}}_{\rm B}^{\rm L}(0) = (S/2L_{\rm D})\exp(-|Y|/L_{\rm D})(-\hat{\boldsymbol{x}})$$
(19)

364 and

$$\boldsymbol{R} = (\Gamma S/2L_{\rm D}) \exp(-|Y|/L_{\rm D})(+\boldsymbol{\hat{y}})$$
(20)

with  $\Gamma$  evaluated at the edge of the core and with vertical averaging understood in (13). Notice that the signs have reverted to those in problem (i). April 6, 2020

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## 367 6. Cross-checks from refraction calculations

The foregoing recoil formulae were rederived in M19 via a completely independent route, 368 in two stages. Firstly, it was shown that an appropriate 'impulse-pseudomomentum theorem' 369 holds for any vortical flow and any wave field with irrotational sources and sinks, provided that 370 the mean flows comply with either (10) or (11). This means that the pseudomomentum rule is 371 guaranteed to hold in any such situation. Then, secondly, explicit linear-theoretic calculations 372 of wave refraction were carried out to yield the incoming and outgoing pseudomomentum 373 fluxes in our three problems. They were found to be precisely consistent with the foregoing 374 results (14), (17), (18), and (20). 375

The calculations used ray theory in all three problems and, in addition, in problem (ii), used the long-wavelength asymptotics and Fresnel-diffractive wake structure analysed in FLS99. FLS99's careful analysis completes the picture described by (2) in an elegant way when both  $k_0 r_0$  and  $\epsilon$  are small, making the Aharonov–Bohm phase jump  $2\pi\alpha$  even smaller, being the product of *two* small quantities as seen in (3).

In problem (ii), the two contributions to R in (18) can be attributed separately to the 381 two refraction effects seen in figure 2. The first contribution in (18), which is the same as 382 (17), arises solely from the wake and the Aharonov–Bohm phase change across it, while 383 the second contribution arises solely from the other refraction effect noted earlier, the  $O(\epsilon)$ 384 rotation of wavecrests in the larger domain outside the wake. This second contribution becomes 385 significant in problem (ii) when the width W of the wavetrain is large enough, accounting for 386 the dependence on W/L. These attributions hold good not only when  $k_0 r_0 \ll 1$  but also 387 when  $k_0 r_0 \gg 1$ , allowing the use of ray theory. In that case the wake has a different structure 388 involving a ray caustic, but continues to account solely for the first contribution in (18) while 389 the  $O(\epsilon)$  rotations of wavecrests outside the wake, which as noted earlier are consistent with 390 ray theory, account solely for the second contribution. In particular, therefore, the Aharonov– 391 Bohm effect is the only relevant refraction effect – as often assumed in the quantum vortex 392 literature – only in the special limiting case for which  $\lim (W/L) = 0$ . 393

In problem (iii), ray theory is used. Because the vortex now has quasigeostrophic structure with finite scale  $L_{\rm D}$  it is the potential vorticity, not the vorticity  $\omega$ , that vanishes outside the vortex core. Therefore the curl-curvature formula of Dysthe (2001) implies that rays passing outside the vortex core, as in figure 1(a), are now deflected at leading order in  $\epsilon$ . The formula tells us that the ray curvature is just  $\omega/C$  where  $C = \frac{1}{2}(g/k)^{1/2}$ , the intrinsic group velocity for deep-water gravity waves. When the formula is used and the calculations carried out (M19 section 8), the result is found to agree perfectly with (20).

# 401 7. Concluding remarks

All the foregoing results depend on the restriction to leading order in  $\epsilon$ , which is essential 402 to the derivation of (10) and (11) and therefore essential to our computations of Bretherton 403 flows, and to the proof of the impulse-pseudomomentum theorem. Yet cases are known in 404 which the pseudomomentum rule holds at higher orders of accuracy in  $\epsilon$ . One of them is the 405 case of figure 3, which requires one further order of accuracy, and in which the mean-flow 406 equations are less simple than (10) and (11). Indeed, in BM03 section 5.2 it was shown that 407 there is a version of that case for which the rule holds to all orders in  $\epsilon$ . On the other hand, 408 as recalled in M19, exceptions to the rule have long been known. 409

There is a major unresolved puzzle here, and a challenge for future work. For now, one may speculate that the present restriction to leading order in  $\epsilon$  may be a limitation of the Kelvin impulse concept, rather than of the pseudomomentum rule itself. Some further discussion of this issue is given in section 9 of M19. 12

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