

# Chapter 4

## Energy and Stability

### 4.1 Energy in 1D

Consider a particle in 1D at position  $x(t)$ , subject to a force  $F(x)$ , so that  $m\ddot{x} = F(x)$ . Define the *kinetic energy* to be

$$T = \frac{1}{2}m\dot{x}^2$$

and the *potential energy* to be

$$V(x) = -\int_a^x F(x') \, dx'$$

where  $a$  is an arbitrary constant. A different choice  $\hat{a}$  would lead to a different potential energy

$$\begin{aligned}\hat{V}(x) &= -\int_{\hat{a}}^x F(x') \, dx' \\ &= -\int_{\hat{a}}^a F(x') \, dx' + V(x),\end{aligned}$$

in effect just adding a fixed constant on to  $V$ . We often choose  $a = 0$  or  $\infty$ .

Note that  $V'(x) = -F(x)$ , and that this is true regardless of the choice of  $a$ .

The dimensions of energy are  $ML^2T^{-2}$  (either from  $[T] = M(LT^{-1})^2$ , or from  $[V] = [\text{force}] \times L = (MLT^{-2})L$ ). In the SI system, it is measured in Joules ( $1 \text{ J} \equiv 1 \text{ kg m}^2/\text{s}^2$ ).

**Examples:** a particle moving vertically under gravity has  $m\ddot{z} = -mg$ , so the potential energy is  $V(z) = -\int_a^z (-mg) \, dz' = mgz + \text{const.}$ . A particle attached to a spring with spring constant  $k$  has  $F = -kx$ , so  $V(x) = -\int_a^x (-kx') \, dx' = \frac{1}{2}kx^2 + \text{const.}$

Multiply the equation of motion,  $F = m\ddot{x}$ , by  $\dot{x}$ :

$$\begin{aligned}
 m\dot{x}\ddot{x} &= \dot{x}F(x) \\
 &= -\dot{x}V'(x) \\
 &= -\frac{dx}{dt} \frac{dV}{dx} \\
 \implies \frac{d}{dt}(\tfrac{1}{2}m\dot{x}^2) &= -\frac{dV}{dt} \\
 \implies T + V &= \text{const.}
 \end{aligned}$$

Hence the total energy  $E = T + V$  is conserved.

Note that if the forcing  $F$  depends on anything other than the position,  $x$ , then energy may not be conserved. For instance, in the damped simple harmonic motion example of §2.2, where a particle attached to a spring was subject to linear friction, we had

$$m\ddot{x} = -kx - c\dot{x}.$$

The potential energy resulting from the spring is  $\frac{1}{2}kx^2$  (+ const.), as above, but it is not possible to define a potential energy for the damping term  $c\dot{x}$ . The total energy (kinetic and potential) combined is therefore  $\frac{1}{2}m\dot{x}^2 + \frac{1}{2}kx^2$ , but

$$\begin{aligned}
 \frac{d}{dt}(\tfrac{1}{2}m\dot{x}^2 + \tfrac{1}{2}kx^2) &= m\dot{x}\ddot{x} + kx\dot{x} \\
 &= \dot{x}(m\ddot{x} + kx) \\
 &= -c\dot{x}^2 \leq 0.
 \end{aligned}$$

The damping causes dissipation of energy.

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When the potential energy is proportional to the mass of the particle, as is the case for instance in vertical motion under gravity where  $V(z) = mgz + \text{const.}$ , it is sometimes useful to consider the potential energy *per unit mass*, which would here be just  $V_u(z) = gz + \text{const.}$  The force on a particle of mass  $m$  is then given by  $F(z) = -mV'_u(z)$ . This can be a useful definition because it allows us to specify the gravitational field without having to know in advance the mass of the particle on which gravity will be acting. However, a major drawback is that there is, unfortunately, no consistent naming convention for “potential energy per unit mass”: some mathematicians call it the “potential field”, others just the “potential”, even though yet others use the word “potential” as a shorthand for “potential energy”. The potential confusion is great.

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## Using the Energy Integral

From

$$\tfrac{1}{2}m\dot{x}^2 + V(x) = E,$$

the “energy integral”, we deduce

$$\dot{x} = \pm \sqrt{\frac{2(E - V)}{m}}.$$

(We have to decide the sign on a case-by-case basis using physical considerations.) So

$$\begin{aligned} \frac{dt}{dx} &= \pm \sqrt{\frac{m}{2(E - V)}} \\ \Rightarrow \quad t &= \int \frac{dt}{dx} dx = \pm \int \sqrt{\frac{m}{2(E - V)}} dx. \end{aligned}$$

**Example:** what is the period of *finite* (i.e., non-infinitesimal) oscillations of a simple pendulum? Suppose that we release a pendulum bob from rest at  $\theta = \theta_0$ . Multiply (2.4),

$$ml\ddot{\theta} = -mg \sin \theta,$$

by  $\dot{\theta}/m$  and integrate:

$$\begin{aligned} l\dot{\theta}\ddot{\theta} &= -g\dot{\theta} \sin \theta \\ \Rightarrow \quad \frac{1}{2}l\dot{\theta}^2 &= g(\cos \theta - \cos \theta_0) \\ \Rightarrow \quad \dot{\theta} &= \pm \sqrt{\frac{2g}{l}(\cos \theta - \cos \theta_0)}. \end{aligned}$$

To determine the period  $\tau$  we consider a quarter-period in which  $\theta$  increases from 0 to  $\theta_0$ :

$$\int_0^{\frac{1}{4}\tau} dt = \int_0^{\theta_0} \frac{d\theta}{\dot{\theta}} \quad \Rightarrow \quad \tau = \sqrt{\frac{8l}{g}} \int_0^{\theta_0} \frac{d\theta}{\sqrt{\cos \theta - \cos \theta_0}}.$$

This is a “complete elliptic integral of the first kind”.

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In fact,  $\tau = 4\sqrt{l/g} K(\sin \frac{1}{2}\theta_0)$  where  $K$  is the appropriate elliptic function.

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## 4.2 Equilibria and Small Oscillations in 1D

A system has an *equilibrium* at  $x = x_0$  if  $F(x_0) = 0$ , because then  $x(t) = x_0 \forall t$  is a solution of the equation of motion. Hence the equilibrium points are the *critical points* of  $V$ , i.e., where  $V'(x_0) = 0$ .

Consider a disturbance around an equilibrium point,  $x = x_0 + \xi(t)$ , where  $\xi$  is initially small. Then

$$\begin{aligned} m\ddot{x} &= F(x) = -V'(x_0 + \xi) \\ \Rightarrow \quad m\ddot{\xi} &= -V'(x_0) - \xi V''(x_0) + O(\xi^2), \end{aligned} \quad (\text{Taylor series})$$

so that

$$\ddot{\xi} + \frac{V''(x_0)}{m}\xi = 0$$

to first order.

If  $V''(x_0) < 0$  then the solution for  $\xi$  is

$$\xi = Ae^{\alpha t} + Be^{-\alpha t}$$

where  $\alpha = \sqrt{-V''(x_0)/m}$ . Hence  $\xi$  grows larger as  $t$  increases (unless  $A = 0$ , which is only true for very special initial conditions), and so the solution does not necessarily stay in a small neighbourhood of  $x_0$ . Such a point is called an *unstable equilibrium*.

If  $V''(x_0) > 0$  then instead  $\xi$  executes SHM with angular frequency

$$\omega = \sqrt{\frac{V''(x_0)}{m}};$$

hence  $\xi$  stays small and the solution stays in a small neighbourhood of  $x_0$ . Such a point is a *stable equilibrium*.

Hence a local maximum of  $V$  is an unstable point, while a minimum is stable. (A point at which  $V''(x_0) = 0$  would require more detailed investigation, taking more terms in the Taylor series.)

**Example:** a particle on a spring. Here  $V(x) = \frac{1}{2}kx^2$ , so since  $V'(0) = 0$  and  $V''(0) = k > 0$ , there is a stable equilibrium at  $x = 0$ . The frequency of oscillations is  $\omega = \sqrt{k/m}$  and the period is  $2\pi\sqrt{m/k}$ , as previously found in §2.2.

### 4.3 Shape of the Potential Energy Function

Consider a graph of  $V(x)$  as shown. We know that  $T + V = E$ , so  $V = E - \frac{1}{2}m\dot{x}^2 \leq E$ . Hence the particle is restricted to regions where  $V(x) \leq E$ .

Consider a particle at rest at the equilibrium point  $x_0$  shown, where  $V$  has a local minimum; then  $E = E_0 \equiv V(x_0)$ . Suppose that the particle is now given a small disturbance. Such a disturbance will change the value of  $E$ , say to  $E_1$  as shown. Then the particle is able to move, but is restricted to the region  $[x_1, x_2]$ ; i.e., it must remain in a neighbourhood of  $x_0$ . This is an alternate way of showing that local minima of  $V$  are stable.

Now suppose that the particle is at another equilibrium point  $x'_0$  which is a local maximum of  $V$ . A small change in  $E$ , to  $E'_1$  say, does *not* restrict the motion of the particle to points near  $x'_0$ . So this is an unstable equilibrium.

Suppose that the particle is at  $x_0$ . What initial speed  $v$  ( $> 0$ ) must we impart to it if it is to travel towards  $+\infty$  and keep going? We need to ensure that  $E > E'_0 \equiv V(x'_0)$ , because otherwise the particle will not be able to reach  $x'_0$ ; so

$$\frac{1}{2}mv^2 + V(x_0) > V(x'_0) \quad \implies \quad v > \sqrt{\frac{2(V(x'_0) - V(x_0))}{m}}.$$

Once it has reached  $x'_0$  it can keep going for ever.

## 4.4 Energy in 3D

### Work

In 3D, the kinetic energy of a particle is given by

$$T = \frac{1}{2}m|\dot{\mathbf{x}}|^2 = \frac{1}{2}m\dot{\mathbf{x}} \cdot \dot{\mathbf{x}}.$$

A force  $\mathbf{F}$  acting on a particle which moves through  $\delta\mathbf{x}$  is said to do *work*  $\delta W = \mathbf{F} \cdot \delta\mathbf{x}$ . The total work done by the force on a particle which moves from  $A$  to  $B$  is

$$W = \int_A^B \mathbf{F} \cdot d\mathbf{x}$$

where the integral follows the path taken by the particle. It is obvious that in general  $W$  depends on the path taken.

The *power* is the rate of doing work, i.e.,  $P = \dot{W}$ . In a time interval  $\delta t$ ,

$$P = \frac{\delta W}{\delta t} = \frac{\mathbf{F} \cdot \delta\mathbf{x}}{\delta t} \quad \implies \quad P = \mathbf{F} \cdot \dot{\mathbf{x}}$$

(in the limit  $\delta t \rightarrow 0$ ). Note that from  $\mathcal{NII}$ ,

$$\begin{aligned} \mathbf{F} &= m\ddot{\mathbf{x}} \\ \implies \mathbf{F} \cdot \dot{\mathbf{x}} &= m\dot{\mathbf{x}} \cdot \ddot{\mathbf{x}} \\ &= \frac{d}{dt}(\frac{1}{2}m\dot{\mathbf{x}} \cdot \dot{\mathbf{x}}) \\ &= \frac{dT}{dt}, \end{aligned}$$

so the power is the rate at which kinetic energy increases.

Forces which are normal to a particle's path do no work: this is because a particle's velocity  $\dot{\mathbf{x}}$  is (by definition) tangential to its path, so  $\mathbf{F} \cdot \dot{\mathbf{x}} = 0$  and hence  $P = 0$ , i.e., the rate of doing work vanishes. For example, a magnetic field does no work on a charged particle, because  $\mathbf{F} = q\mathbf{v} \times \mathbf{B}$  is perpendicular to  $\mathbf{v}$ , and the field therefore neither increases nor decreases the particle's kinetic energy. Similarly, the tension in the string of a simple pendulum does no work on the bob.

## Conservative Forces

A force field  $\mathbf{F}(\mathbf{x})$  is said to be *conservative* iff the work done by the force on a particle moving from any point  $A$  to any other point  $B$  is independent of the path taken: i.e., iff  $\int_A^B \mathbf{F} \cdot d\mathbf{x}$  is path-independent. We know from the *Vector Calculus* course that this is equivalent to saying that  $\mathbf{F}$  is conservative iff

$$\boxed{\mathbf{F} = -\nabla V}$$

for some function  $V(\mathbf{x})$ , called the potential energy. (This is the 3D equivalent of  $F = -V'(x)$  in 1D; in 1D *all* force fields  $F(x)$  are conservative.)

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But in 3D, not all force fields are conservative, because (as shown in the *Vector Calculus* course) the value of a line integral  $\int_A^B \mathbf{F} \cdot d\mathbf{x}$  depends, in general, on the path taken.

What is the work done in moving a particle from a starting point  $A$  round a closed path back to the starting point? For a conservative force, the answer must be path-independent and must therefore be zero. But for a non-conservative force this does not apply and the force may have to do work just to get the particle back to where it started. This non-zero work done is generally dissipated, for example as heat.

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For a conservative force, the work done is

$$W = \int_A^B \mathbf{F} \cdot d\mathbf{x} = - \int_A^B \nabla V \cdot d\mathbf{x} = - \int_A^B dV = -[V(\mathbf{x})]_A^B,$$

i.e., equal to the decrease in potential energy. We can also prove that the total energy  $E = T + V$  is conserved:

$$\begin{aligned} \frac{dE}{dt} &= \frac{d}{dt} \left( \frac{1}{2} m \dot{\mathbf{x}} \cdot \dot{\mathbf{x}} \right) + \frac{d}{dt} V(\mathbf{x}) \\ &= m \dot{\mathbf{x}} \cdot \ddot{\mathbf{x}} + \frac{\partial V}{\partial x} \frac{dx}{dt} + \frac{\partial V}{\partial y} \frac{dy}{dt} + \frac{\partial V}{\partial z} \frac{dz}{dt} \\ &= \dot{\mathbf{x}} \cdot \mathbf{F} + \nabla V \cdot \dot{\mathbf{x}} \\ &= 0. \end{aligned}$$

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Note that, just as in 1D, a local minimum of  $V(\mathbf{x})$  is a stable equilibrium, and a local maximum is unstable — we can see this by considering the shape of  $V(\mathbf{x})$  as in §4.3. (Quite easy to do in 2D; but almost impossible in 3D!) A saddle point of  $V$  is also unstable because the particle can move “downhill” from the saddle.

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## Examples of Potential Energy Functions in 3D

The force due to gravity on a particle of mass  $m$  near the Earth’s surface is  $\mathbf{F} = m\mathbf{g}$ . But

$$\mathbf{g} = \begin{pmatrix} 0 \\ 0 \\ -g \end{pmatrix} = \nabla(-gz),$$

so  $V(\mathbf{x}) = mgz + \text{const.}$ : exactly the same answer as for 1D vertical motion.

A spring with spring constant  $k$  and natural length  $l$  attached to a fixed point  $O$ , but otherwise free to move in any direction in 3D, exerts a force

$$\mathbf{F}(\mathbf{r}) = -k(r - l)\hat{\mathbf{e}}_r$$

towards  $O$ , where  $r = |\mathbf{r}|$  and  $\hat{\mathbf{e}}_r = \mathbf{r}/r$  is the unit radial vector. We note that

$$\nabla\left\{\frac{1}{2}(r - l)^2\right\} = (r - l)\hat{\mathbf{e}}_r$$

so

$$V(\mathbf{r}) = \frac{1}{2}k(r - l)^2 + \text{const.},$$

the same as the 1D potential energy  $\frac{1}{2}kx^2$  (+ const.) where  $x$  is the extension.

A uniform electric field  $\mathbf{E}$  acting on a charge  $q$  produces a force  $q\mathbf{E}$ . But

$$\nabla(\mathbf{E} \cdot \mathbf{x}) = \nabla(E_1x_1 + E_2x_2 + E_3x_3) = (E_1, E_2, E_3)^T = \mathbf{E},$$

so the potential energy is  $-q\mathbf{E} \cdot \mathbf{x}$ .

The gravitational force on a particle of mass  $m_2$  with position vector  $\mathbf{r}_2$  due to a particle of mass  $m_1$  at  $\mathbf{r}_1$  is

$$\mathbf{F} = -\frac{Gm_1m_2}{|\mathbf{r}|^2}\hat{\mathbf{e}}_r$$

from §1.5.1, where  $\mathbf{r} = \mathbf{r}_2 - \mathbf{r}_1$  is the relative position vector and  $\hat{\mathbf{e}}_r = \mathbf{r}/|\mathbf{r}|$ . So

$$V(\mathbf{r}) = -\frac{Gm_1m_2}{|\mathbf{r}|} \tag{4.1}$$

(where we choose the arbitrary constant so that  $V = 0$  at infinity). In particular, the gravitational potential energy produced by the Earth (a mass  $M$  at the origin) acting on a particle of mass  $m$  at  $\mathbf{r}$  is

$$V(\mathbf{r}) = -\frac{GMm}{r} \tag{4.2}$$

where  $r = |\mathbf{r}|$ .

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In (4.1) we can consider the potential  $V$  to be a function of two variables  $\mathbf{r}_1, \mathbf{r}_2$ :

$$V(\mathbf{r}_1, \mathbf{r}_2) = -\frac{Gm_1m_2}{|\mathbf{r}_2 - \mathbf{r}_1|}.$$

The force on the second particle due to the first is then given by  $-\nabla_2 V$  where  $\nabla_2$  denotes the gradient operator taken with respect to  $\mathbf{r}_2$  (keeping  $\mathbf{r}_1$  fixed), i.e.,

$$\nabla_2 \equiv \begin{pmatrix} \frac{\partial}{\partial(\mathbf{r}_2)_1} \\ \frac{\partial}{\partial(\mathbf{r}_2)_2} \\ \frac{\partial}{\partial(\mathbf{r}_2)_3} \end{pmatrix}.$$

The *same* potential function also gives us the force on the first particle due to the second, which is  $-\nabla_1 V$ ; the symmetry in  $\mathbf{r}_1$  and  $\mathbf{r}_2$  ensures that the forces are equal and opposite. This idea can be extended to a system of  $n$  particles, with a potential function  $V(\mathbf{r}_1, \dots, \mathbf{r}_n)$  that depends on all of the interparticle distances  $r_{ij} = |\mathbf{r}_i - \mathbf{r}_j|$ . The total force on the  $i^{\text{th}}$  particle due to the others is then given by  $-\nabla_i V$ .

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## 4.5 Escape Velocity

The *escape velocity* is the minimum initial speed that would need to be imparted to a particle in a gravitational field to enable it to get arbitrarily far away.

For example, consider the potential energy (4.2) for a particle moving in the Earth's gravitational field,  $-GMm/r$ , as shown in the diagram. If the particle's total energy is  $E_1 < 0$ , then it is restricted to

$$r \leq r_1 = \frac{GMm}{-E_1};$$

if its total energy is instead  $E_2 \geq 0$  then its motion is unrestricted and the particle can escape to  $\infty$ .

If the particle starts from  $r = r_0$  with speed  $v$  then

$$E = \frac{1}{2}mv^2 - \frac{GMm}{r_0}.$$

The escape velocity, i.e., the minimum value of  $v$  required to ensure that  $E \geq 0$ , is

$$v_{\text{escape}} = \sqrt{\frac{2GM}{r_0}}.$$

If a space ship starts on the surface of the Earth at  $r = R$ , then using  $g = GM/R^2$  (from §1.5.1) we obtain an escape velocity of

$$\sqrt{2gR} \approx 11.2 \text{ km/s}$$

required to clear the Earth's gravitational field.



## 4.6 Motion with One Degree of Freedom

A particle may follow a trajectory in three dimensions which can be described by a single parameter  $q(t)$ . For example, the location of a bead moving along a wire can either be described by its position vector  $\mathbf{r}$  in 3D, or instead by the scalar variable  $s$  defined as the arc-length measured along the wire from a fixed point. So long as we know the shape of the wire, the single variable  $s$  tells us everything we need to know about the bead's position. Such a system is known as having *one degree of freedom*, and we can treat it as being effectively one-dimensional.

For such a system, described by a parameter  $q(t)$ , we can apply all the methods of §§4.1–4.3 for motion in 1D, but we first need to obtain an equation of motion in the form

$$m\ddot{q} = F^*(q)$$

for some function  $F^*(q)$  (which is not necessarily the actual force). Then, just as in §4.1, we can define a (pseudo-)kinetic energy  $T^* = \frac{1}{2}m\dot{q}^2$  and a (pseudo-)potential energy  $V^* = -\int_a^q F^*(q') dq'$ , and deduce that

$$T^* + V^* = E^*$$

is constant. Note however that  $T^*$  and  $V^*$  may well not be equal to the true kinetic and potential energies of the system.

For example, consider a simple pendulum swinging in a plane. This system has one degree of freedom, because the bob's position in 3D is specified completely by  $\theta(t)$ . We therefore have a choice of approaches:

- Treat the pendulum as a system with one degree of freedom. Starting from the equation of motion (2.4) in the appropriate form,

$$m\ddot{\theta} = -\frac{mg}{l} \sin \theta \equiv F^*(\theta),$$

define  $T^* = \frac{1}{2}m\dot{\theta}^2$  and

$$V^* = -\int_a^\theta F^*(\theta') d\theta' = \frac{mg}{l} \int_a^\theta \sin \theta' d\theta' = -\frac{mg}{l} \cos \theta + \text{const.}$$

Ignoring the arbitrary constant, we therefore have that

$$E^* = T^* + V^* = \frac{1}{2}m\dot{\theta}^2 - \frac{mg}{l} \cos \theta$$

is constant. This approach is most useful when we are able to write down the equation of motion straight away but we do not know the true energy  $E$  of the system; it allows us to find an conserved quantity (namely  $E^*$ ). Small oscillations can be investigated using the results of §4.2 (including the formula for the frequency,  $\omega = \sqrt{V^{*''}(q_0)/m}$ ) directly.

- Use the full 3D system. The speed of the bob is  $l\dot{\theta}$ , so the (true) kinetic energy of the system is  $T = \frac{1}{2}ml^2\dot{\theta}^2$ . The bob is at a height  $l \cos \theta$  below the point of suspension  $O$ , so the (true) potential energy is  $V = -mgl \cos \theta$ . We therefore deduce that

$$E = \frac{1}{2}ml^2\dot{\theta}^2 - mgl \cos \theta$$

is constant. To obtain the equation of motion, differentiate with respect to time:

$$0 = ml^2\dot{\theta}\ddot{\theta} + mgl\dot{\theta} \sin \theta \quad \implies \quad \ddot{\theta} = -\frac{g}{l} \sin \theta.$$

This approach is most useful when we *cannot* initially write down the equation of motion of the system, but we *can* calculate its energy instead; the steps above then lead us to the equation of motion. To calculate the frequency of small oscillations about stable equilibria, it is necessary to consider a small disturbance and expand the equation of motion using Taylor Series as in §4.2: the formula given there for the frequency cannot be applied directly.

These two approaches are entirely consistent, because  $E$  and  $E^*$  differ only by a constant factor:  $E^* = E/l^2$ .

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A general system with one degree of freedom has  $\mathbf{x} = \mathbf{x}(q)$ , so that

$$\dot{\mathbf{x}} = \frac{d\mathbf{x}}{dq} \frac{dq}{dt} = \mathbf{x}' \dot{q}$$

where a prime denotes differentiation with respect to  $q$ . The (true) kinetic energy is therefore

$$T = \frac{1}{2}m|\dot{\mathbf{x}}|^2 = \frac{1}{2}m|\mathbf{x}'|^2\dot{q}^2 = |\mathbf{x}'|^2T^*.$$

Thus the factor relating  $T$  to  $T^*$  is  $|\mathbf{x}'|^2$ .

In the case of a pendulum, it is obvious that  $|\mathbf{x}'| = |d\mathbf{x}/d\theta| = l$ , and the factor relating  $T$  to  $T^*$  is  $l^2$  as found above.

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