

1 Constraints

(a) Reproduce the constraint equations by varying the action

$$S = \int d^4x \sqrt{h} N \left\{ \frac{M_{\text{Pl}}^2}{2} \left[{}^{(3)}R + K_{ij} K^{ij} - K^2 \right] + P(X, \phi) \right\}.$$

with respect to N and N^i .

(b) Solve the $\delta S / \delta N^i$ constraint to find δN , working in flat gauge to linear order.

2 Gauge Transformations

(a) Derive the linear-order gauge transformations of A , B , ψ and h_{00} , for a generic change of coordinates $x^\mu \rightarrow x^\mu + \epsilon^\mu$.

(b) Derive the gauge transformation from Newtonian gauge to flat gauge and viceversa. In particular, given some generic perturbations $\{A^N, h_{00}^N, \varphi^N\}$ in Newtonian gauge, determine the corresponding perturbations $\{h_{00}^f, \psi^f, \varphi^f\}$ in flat gauge.

3 Conservation of \mathcal{R}

In the lecture, we prove the conservation of \mathcal{R} and γ_{ij} on superHubble scales in the presence of a generic energy-momentum tensor $T_{\mu\nu}$ by working in comoving gauge. Prove again the conservation of \mathcal{R} by working in Newtonian gauge. In particular, you might want to start with the change of coordinates

$$\epsilon^\mu = \left\{ \epsilon(t), \lambda x^i \right\}.$$

and show that the gauge transformations are

$$\begin{aligned} \Phi &= -\dot{\epsilon}, & \Psi &= H\epsilon - \frac{\lambda}{3}, \\ \delta\rho &= -\dot{\rho}\epsilon, & \delta u &= \epsilon, & \pi^S &= 0, \\ \delta p &= -\dot{p}\epsilon, & \varphi &= -\epsilon\dot{\phi}. \end{aligned}$$

Then use the scalar part of the ij components of the Einstein's equation,

$$k_i k_j (\Phi - \Psi) = 0,$$

to impose the physicality condition on $\epsilon(t)$. Your final result should be

$$\mathcal{R} = \frac{\lambda}{3}, \quad \varphi = -\dot{\phi} \frac{\mathcal{R}}{a} \int_T^t a(t') dt', \quad \Phi = \Psi = \mathcal{R} \left[-1 + \frac{H}{a} \int_T^t a(t') dt' \right].$$

4 Power Spectrum and Correlation Function

(a) Writing the Fourier transform of $\delta(\mathbf{k})$ and $\delta(\mathbf{k}')$, derive the expression for $\langle \delta(\mathbf{k}) \delta(\mathbf{k}') \rangle$ in terms of the correlation function $\xi(r)$.

(b) Writing the Fourier transform of $\delta(\mathbf{x})$ and $\delta(\mathbf{x}' + \mathbf{r})$, derive the expression for $\langle \delta(\mathbf{x}) \delta(\mathbf{x}' + \mathbf{r}) \rangle$ in terms of the power spectrum $P(k)$.

(c) The power spectrum and correlation function are related to each other through spherical Bessel transformations $P(k) = 4\pi \int dr r^2 \xi(r) j_0(kr)$ and $\xi(r) = 1/(2\pi^2) \int dk k^2 P(k) j_0(kr)$. Show that closure of this relation, i.e., $P \rightarrow \xi \rightarrow P$ yields the closure relation for spherical Bessel functions

$$\int_0^\infty dx x^2 j_\alpha(ux) j_\alpha(vx) = \frac{\pi}{2u^2} \delta^{(D)}(u - v).$$

5 Gaussian PDF and Edgeworth expansion

Consider the d_y dimensional random vector \mathbf{y} described by the multivariate Gaussian PDF

$$\mathbb{P}_G(\mathbf{y}) = \frac{1}{\sqrt{(2\pi)^{d_y} |M|}} \exp \left[-\frac{1}{2} \mathbf{y}^T M^{-1} \mathbf{y} \right].$$

Now split the random vector \mathbf{y} into two components $\mathbf{y} = (\mathbf{m}, \mathbf{n})$ of dimensions d_m and d_n with covariance matrix

$$M = \begin{pmatrix} \langle m, m \rangle & \langle m, n \rangle \\ \langle n, m \rangle & \langle n, n \rangle \end{pmatrix} = \begin{pmatrix} A & B \\ B^T & D \end{pmatrix} = \begin{pmatrix} S & T \\ T^T & V \end{pmatrix}^{-1}.$$

(a) *Marginalization and conditional probability* – By explicitly integrating out the variables n , show that

$$\mathbb{P}_G(\mathbf{m}) = \frac{1}{\sqrt{(2\pi)^{d_m} |A|}} \exp \left[-\frac{1}{2} \mathbf{m}^T A^{-1} \mathbf{m} \right].$$

It will be useful to use the formulae for block inversion of the matrix

$$M^{-1} = \begin{pmatrix} S & T \\ T^T & V \end{pmatrix} = \begin{pmatrix} A^{-1} + A^{-1}B(D - B^T A^{-1}B)^{-1}B^T A^{-1} & -A^{-1}B(D - B^T A^{-1}B)^{-1} \\ -(D - B^T A^{-1}B)^{-1}B^T A^{-1} & (D - B^T A^{-1}B)^{-1} \end{pmatrix}$$

and to complete the squares. For the determinant one has $|M| = |A|/|V|$ with $V = (D - B^T A^{-1}B)^{-1}$. Now show that the conditional PDF of variables n given m can be expressed as

$$\mathbb{P}_G(\mathbf{n}|\mathbf{m}) = \frac{1}{\sqrt{(2\pi)^{d_n} |\Sigma|}} \exp \left[-\frac{1}{2} (\mathbf{n} - \boldsymbol{\mu})^T \Sigma^{-1} (\mathbf{n} - \boldsymbol{\mu}) \right],$$

where $\boldsymbol{\mu} = B^T A^{-1} \mathbf{m}$ and $\Sigma = V^{-1} = D - B^T A^{-1} B$.

(b) *Edgeworth expansion* – Similar to what is discussed in the lecture notes for the one-point PDF, the Edgeworth expansion of a Gaussian random field with three point function $\mathcal{R}_{ijl} \ll 1$ can be expressed as

$$\mathbb{P}(\mathbf{y}) = \left(1 - \frac{\mathcal{R}_{ijl}}{3!} \frac{\partial}{\partial y_i} \frac{\partial}{\partial y_j} \frac{\partial}{\partial y_l} \right) \mathbb{P}_G(\mathbf{y}).$$

Evaluate the PDF explicitly and then simplify the result for the case of a diagonal covariance matrix $C_{ij} = \delta_{ij}^{(K)} C_i$. Show that the characteristic function is given by

$$\mathcal{M}(i\mathbf{J}) = \left(1 - (-i)^3 \frac{\mathcal{R}_{ijl}}{3!} J_i J_j J_l \right) \exp \left[-\frac{1}{2} J_m C_{mn} J_n \right].$$

Use the characteristic function to evaluate the expectation values $\langle f_i \rangle$, $\langle f_i f_j \rangle$ and $\langle f_i f_j f_l \rangle$.

6 Primordial non-Gaussianity

In the local non-Gaussian model, higher order corrections to the linear solution $\mathcal{R}^{(1)}$ arise simply from powers of $\mathcal{R}^{(1)}$, that is, at second-order

$$\mathcal{R}(\mathbf{x}) = \mathcal{R}^{(1)}(\mathbf{x}) + \mathcal{R}^{(2)}(\mathbf{x}) + \dots = \mathcal{R}^{(1)} + \frac{3}{5} f_{\text{NL}} ((\mathcal{R}^{(1)})^2 - \langle (\mathcal{R}^{(1)})^2 \rangle) + \dots,$$

where we assume a Gaussian isotropic model for the linear solution $\mathcal{R}^{(1)}$, that is, its Fourier transform satisfies $\langle \mathcal{R}^{(1)}(\mathbf{k}) \mathcal{R}^{(1)}(\mathbf{k}') \rangle = (2\pi)^3 P_{\mathcal{R}}(k) \delta^{(D)}(\mathbf{k} + \mathbf{k}')$ with power spectrum $P(k)$. Using the convolution theorem, show that the solution for the second-order solution $\mathcal{R}^{(2)}$ can be expressed in the form,

$$\mathcal{R}^{(2)}(\mathbf{k}) = \int \frac{d^3 \mathbf{k}'}{(2\pi)^3} \int \frac{d^3 \mathbf{k}''}{(2\pi)^3} \left[\mathcal{R}^{(1)}(\mathbf{k}') \mathcal{R}^{(1)}(\mathbf{k}'') - \langle \mathcal{R}^{(1)}(\mathbf{k}') \mathcal{R}^{(1)}(\mathbf{k}'') \rangle \right] \delta^{(D)}(\mathbf{k} - \mathbf{k}' - \mathbf{k}'')$$

Hence, find the bispectrum for the local model $\langle \mathcal{R}(\mathbf{k}_1) \mathcal{R}(\mathbf{k}_2) \mathcal{R}(\mathbf{k}_3) \rangle = (2\pi)^3 \delta^{(D)}(\mathbf{k}_1 + \mathbf{k}_2 + \mathbf{k}_3) B(k_1, k_2, k_3)$,

$$B^{\text{loc}}(k_1, k_2, k_3) = \frac{6}{5} f_{\text{NL}} [P_{\mathcal{R}}(k_1)P_{\mathcal{R}}(k_2) + P_{\mathcal{R}}(k_2)P_{\mathcal{R}}(k_3) + P_{\mathcal{R}}(k_3)P_{\mathcal{R}}(k_1)].$$

7 Projections of statistically-homogeneous and isotropic random fields

Consider a random field $F(\mathbf{x})$ that is statistically homogeneous and isotropic and has zero mean. The radial projection of the gradient of this field on the surface of the sphere $|\mathbf{x}| = r$ defines a field $f(\hat{\mathbf{n}}) \equiv \hat{\mathbf{n}} \cdot \nabla F(r\hat{\mathbf{n}})$.

(a) Using the Rayleigh plane-wave expansion,

$$e^{i\mathbf{k} \cdot \mathbf{x}} = 4\pi \sum_{lm} i^l j_l(kr) Y_{lm}^*(\hat{\mathbf{k}}) Y_{lm}(\hat{\mathbf{x}}),$$

where $j_l(kr)$ are spherical Bessel functions, show that the spherical multipoles of $f(\hat{\mathbf{n}})$ are related to the Fourier transform of $F(\mathbf{x})$ by

$$f_{lm} = 4\pi i^l \int \frac{d^3\mathbf{k}}{(2\pi)^3} F(\mathbf{k}) k j_l'(kr) Y_{lm}^*(\hat{\mathbf{k}}),$$

where $j_l'(x) = dj_l/dx$.

(b) Hence show that $f(\hat{\mathbf{n}})$ is statistically isotropic with angular power spectrum

$$C_l = 4\pi \int d \ln k \Delta_F^2(k) k^2 [j_l'(kr)]^2,$$

where $\Delta_F^2(k)$ is the power spectrum of $F(\mathbf{x})$. Sketch $[j_l'(kr)]^2$ and so deduce that the mapping between k and l is less sharp for the projection of the radial gradient of a field compared to the projection of the field itself.